The Level Density of Superheavy Nuclei

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Motivation

The experimental trend of nuclear properties (Q_a -values and half-lives) as well as the production cross sections of SHE reveal an increasing stability of nuclei approaching N=184 and indicate quite large shell effects beyond Z=114.

This is in agreement with the predictions of relativistic and nonrelativistic meanfield models which expect the "island of stability" at

Z=120-126 and N=172 or 184.

P.G. Reinhard, Rep. Prog. Phys. 52 (1989); P. Ring, Prog. Par. Nucl. Phys. 37 (1996); M. Bender et al., Rev. Mod. Phys. 75 (2003); J Meng et al., Prog. Part. Nucl. Phys. 57 (2006); J.J. Li, Phys. Lett. B 732 (2014).

In the (N,Z)-plane, the line, along which all new SHE were discovered in the actinide-based reactions with ⁴⁸Ca beam, just approaches this region.

Yu. Ts. Oganessian, J. Phys. G34 (2007); Phys. Rev. C 87 (2013)

If the shell-effects at Z=120 is strong, then there is a hope to synthesize new SHE elements with $Z \ge 120$ using the present experimental set up with available stable projectiles and targets.

Motivation

The investigation of the nuclear properties and the shell structure of elements with 100 < Z < 130 is important from the perspective of future experiments.

- The level density, as a function of excitation energy, is required to calculate the survival probability and, correspondingly, the production cross section of SHE.

- The phenomenological values of level density parameters have to be verified in the microscopical calculations.

In this work, the modified TCSM is used to investigate the **level density of** 100 < Z < 130 nuclei, which belongs to the alpha-chains of ^{296,298,300}120 nuclei.

Previously, the TCSM was modified to reveal the trends of shell structure, nuclear binding energies, Q_{α} -values, and quasiparticle states of SHE.

A.N. Kuzmina, G.G. Adamian and N.V. Antonenko, W. Scheid, Phys. Rev. C 85 (2012)

Two Center Shell Model (J. Maruhn and W. Greiner, Z. Phys. 251 (1972) 431)



Description of the Ground State

In the present work we choose the following shape parametrization:

the elongation λ = L/2R₀
the case of deformation β = a/b = β₁ = β₂
the neck parameter ε = E₀/E' = 0
the mass asymmetry η = (A₁-A₂)/(A₁+A₂) = 0



Hamiltonian of model

$$H = -\frac{\hbar^2}{2m} \nabla^2 + V(\rho, z) + V_{Is} + V_{I^2},$$

$$V(\rho, z) = \begin{cases} \frac{1}{2}m\omega_z^2(z-z_1)^2 + \frac{1}{2}m\omega_\rho^2\rho^2, & z < z_1, & \text{m-nucleon mass,} \\ \frac{1}{2}m\omega_\rho^2\rho^2, & z_1 < z < z_2, & \omega_\rho/\omega_z = a/b = \beta, \\ \frac{1}{2}m\omega_z^2(z-z_2)^2 + \frac{1}{2}m\omega_\rho^2\rho^2, & z > z_2. & z_1 - z_2 = 2R_0\lambda - 2a, \\ & \omega_\rho = \beta\omega_0R_0/a, \\ & h\omega_0 = 4I \text{ MeV } A^{-1/3} \end{cases}$$

$$V_{Is} = -\frac{2\hbar\kappa}{m\omega_0'} (\nabla V \times \mathbf{p})\mathbf{s}$$

$$h\omega_0' = 4I \text{ MeV } A^{-1/3}$$

$$V_{I^2} = -\kappa\mu\hbar\omega_0' \mathbf{l}^2 + \kappa\mu\hbar\omega_0' \frac{\mathcal{N}(\mathcal{N} + 3)}{2}\delta_{if}$$

Modification:In order to improve the description of the nuclear spins and paritieswe introduce the weak dependence of κ , μ – on (*N-Z*).

Modification of the two center shell model: details

The momentum-dependent part of the Hamiltonian consists of the sl- and l²-like terms with the parameters $\kappa_{n,p}$ and $\mu_{n,p}$, respectively.

$$\kappa_n = -0.076 + 0.0058 (N-Z) - 6.53 \times 10^{-5} (N-Z)^2 + 0.002A^{1/3}$$

 $\mu_n = 1.598 - 0.0295 (N-Z) + 3.036 \times 10^{-4} (N-Z)^2 - 0.095A^{1/3}$

 $\kappa_{\rm p}$ =0.0383 + 0.00137 (N-Z) - 1.22×10⁻⁵ (N-Z)² - 0.003A^{1/3} μ_{p} =0.335 + 0.01 (N-Z) - 9.367×10⁻⁵ (N-Z)² + 0.003A^{1/3}

G.G.Adamian, N.V. Antonenko, and W. Scheid, Phys. Rev. C 81, 024320 (2010) G.G.Adamian, N.V. Antonenko, S.N. Kuklin, and W. Scheid, Phys. Rev. C 82, 054304 (2010)

Results of TCSM Calculations for ³⁰⁰120 chain.

Α	Z	δE_{sh}	λ	β	β_2	β_4
260	100	-4.13	1.18	1.28	0.257	0.025
264	102	-3.93	1.18	1.28	0.257	0.025
268	104	-3.42	1.16	1.22	0.239	0.014
272	106	-4.29	1.14	1.08	0.245	-0.017
276	108	-5.35	1.12	1.00	0.235	-0.036
280	110	-5.40	1.10	0.96	0.211	-0.044
284	112	-5.40	1.08	0.90	0.193	-0.059
288	114	-6.92	1.06	0.86	0.168	-0.067
292	116	-6.50	0.94	0.90	-0.095	-0.001
296	118	-7.50	1.02	0.94	0.058	-0.028
300	120	-9.20	1.00	0.88	0.036	-0.040
304	122	-9.10	1.00	0.92	0.023	-0.026
308	124	-11.30	0.98	0.90	-0.012	-0.023
312	126	-9.98	0.98	0.90	-0.012	-0.023
316	128	-6.31	0.98	0.90	-0.012	-0.023
320	130	-4.52	0.98	0.90	-0.012	-0.023

Level Density in the Superfluid Nuclear Model

Nuclear level density:

$$\rho = \frac{\exp[S(\beta, \lambda_Z, \lambda_N)]}{(2\pi)^{3/2}\sqrt{D}}.$$

Entropy:

$$S(\alpha_N, \alpha_Z, \beta) = 2 \sum_{k=Z,N} \sum_{\nu} \left\{ \ln\left[1 + \exp(-\beta E_{\nu k})\right] + \frac{\beta E_{\nu k}}{1 + \exp(\beta E_{\nu k})} \right\}$$

where:

$$D = \begin{vmatrix} \frac{\partial^2 S}{\partial \beta^2} & \frac{\partial^2 S}{\partial \beta \partial \mu_Z} & \frac{\partial^2 S}{\partial \beta \partial \mu_N} \\ \frac{\partial^2 S}{\partial \beta \partial \mu_Z} & \frac{\partial^2 S}{\partial \mu_Z^2} & 0 \\ \frac{\partial^2 S}{\partial \beta \partial \mu_N} & 0 & \frac{\partial^2 S}{\partial \mu_N^2} \end{vmatrix}$$

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P. Decowski, W. Grochulski, A. Marcinkowski, K. Siwek, and Z. Wilhelmi, Nucl. Phys. A 110 G.D. Adeev and P.A. Cherdantsev, Yadernaya Fizika 21, 491 (1975)

Level Density in the Superfluid Nuclear Model

- *N* number of neutrons,
- *Z* number of protons,
- $T=\beta^{-1}$ temperature,
- $\lambda_{\rm N=} \alpha_{\rm N} / \beta$ -chemical potential for neutrons,
- $\lambda_{Z} = \alpha_{Z} / \beta$ -chemical potential for protons,
- $\epsilon_{vN,} \epsilon_{vZ}$ -neutron, proton single- particle energies,
- G_N, G_Z -pairing constants.

 $E_{\nu k} = \sqrt{(\epsilon_{\nu k} - \lambda_k)^2 + \Delta_k^2} \quad \text{-quasiparticle energies of neutrons } (k=N) \text{ and protons } (k=Z).$ Equations determining the saddle point: $\boxed{N - \sum \left(1 - \frac{\epsilon_{\nu N} - \lambda_Z}{1 - \epsilon_{\nu N} - \lambda_Z} + \frac{\beta E_{\nu N}}{1 - \epsilon_{\nu Z} - \lambda_Z}\right)} = Z - \sum \left(1 - \frac{\epsilon_{\nu Z} - \lambda_Z}{1 - \epsilon_{\nu Z} - \lambda_Z} + \frac{\beta E_{\nu Z}}{1 - \epsilon_{\nu Z} - \lambda_Z}\right)$

$$N = \sum_{\nu} \left(1 - \frac{c_{\nu N} - \lambda_Z}{E_{\nu N}} \tanh \frac{\beta E_{\nu N}}{2} \right), \quad Z = \sum_{\nu} \left(1 - \frac{c_{\nu Z} - \lambda_Z}{E_{\nu Z}} \tanh \frac{\beta E_{\nu Z}}{2} \right)$$
$$\frac{2}{G_N} = \sum_{\nu} \frac{\tanh \left(\beta E_{\nu N}/2\right)}{E_{\nu N}}, \quad \frac{2}{G_Z} = \sum_{\nu} \frac{\tanh \left(\beta E_{\nu Z}/2\right)}{E_{\nu Z}}$$
$$E(T) = \sum_{k=N,Z} \left\{ \sum_{\nu} \epsilon_{\nu k} \left(1 - \frac{\epsilon_{\nu N} - \lambda_Z}{E_{\nu N}} \tanh \frac{\beta E_{\nu N}}{2} \right) - \frac{\Delta_k^2}{G_k} \right\}$$

 $U = E_{Z,N}(T) - E_{Z,N}(0)$

Intrinsic Level Densities in Heaviest Nuclei



Due to the smaller densities of the singleparticle states near the Fermi surfaces in the nuclei with the closed shell in the ground state, the LD is smaller for magic or nearly magic nuclei.

U <10 MeV the LD of Z=120 and Z=104 isotopes are comparable (pairing).

U ~(10-30) MeV,

the LD in Z = 120 isotopes are one order of magnitude smaller than those in Rf (shell effects).

U >30 MeV

the LD of Z=120 and Z=104 isotopes are comparable (shell effects faded).

Fermi Gas Expression for the Level Density

The level density parameter, as a function of excitation energy a(U) is required to calculate the survival probability and, correspondingly, the production cross section of the heavies nuclei.

a(U) is extracted by fitting the calculated level density by the Fermi-Gas expression.

$$\rho_{FG}(U) = \frac{\sqrt{\pi}}{12a^{1/4}U^{5/4}} \exp\left[2\sqrt{aU}\right]$$
$$U = aT^2, \quad S = 2aT = 2\sqrt{aU}$$

For particles moving in spherical potential well of radius $R = r_0 A^{1/3}$, $r_0 = 1.2$ fm, one can estimate:

$$a = \left(\frac{\pi}{3}\right)^{4/3} \frac{2m_N r_0^2}{\hbar^2} A \approx \frac{A}{13.5}$$

In the phenomenological calculations of surviving probabilities one usually takes

a=A/(10-12) MeV⁻¹

Extraction of the Level Density Parameter

We found that the best fit of the calculated LD with the Fermi gas expression is achieved if one uses the level density parameter

$$a(U)=S^2/4U,$$

where S and U are calculated

$$S = 2 \sum_{k=Z,N} \sum_{\nu} \left\{ \ln[1 + \exp(-\beta E_{k\nu})] + \frac{\beta E_{k\nu}}{1 + \exp(\beta E_{k\nu})} \right\},$$
$$E_{Z,N}(T) = \sum_{k=Z,N} \left\{ \sum_{\nu} \varepsilon_{k\nu} \left(1 - \frac{\varepsilon_{k\nu} - \lambda_k}{E_{k\nu}} \tanh \frac{1}{2} \beta E_{k\nu} \right) - \frac{\Delta_k^2}{G_k} \right\}$$
$$U = E_{Z,N}(T) - E_{Z,N}(0).$$

Note, that one does not need to make a back-shift of excitation energy as in the phenomenological treatment.



Fit of the calculated level density with the Fermi-Gas Expression.



Dependence of Level Density Parameter on Shell Corrections





-For Z < 116, one can use the parametrization $a \cong A/(12-14) MeV^{-1}$

-For Z > 116, the better fit of the calculated values is $a \cong A/(14-17)MeV^{-1}$.

Level Density Parameter and Shell Corrections

*With increase of excitation energy, the correlation between level density parameter *a* and shell corrections δE_{sh} is destroyed and the dependence of *a* parameter on mass number becomes rather smooth.

Shared on the study of the dependencies of **a** on E_{sh} and U, one can use the following parametrization of the level density parameter:

A.B. Ignatyuk, G.N. Smirenkin, and A.S. Tishin, Yadernaya Fizika 21, 485(1975)

$$\alpha = 0.118 \text{ MeV}^{-1}, \, \beta = -0.53 \times 10^{-4} \text{ MeV}^{-1}, \, E'_D = 27 \text{ MeV}$$

Level Density Parameter as a function of N and Z



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Summary

• The intrinsic level densities have been microscopically calculated in the region of superheavy nuclei. The method of calculations have been tested for light nuclei with known experimental data.

• **The level density parameters**, which are used in the Fermi-gas model, have been determined for the nuclei of alpha-decay chains containing ²⁹⁶120,²⁹⁸120,³⁰⁰120.

• The dependencies of the level density parameter on the shell correction and excitation energy have been studied. **The damping factor** in the dependence of the level density parameter on excitation energy is found to be

$$E'_D = 27 \text{ M} \Rightarrow B$$

• For the superheavy nuclei considered, the level density parameters are approximately A/(11-13) MeV at excitation energies corresponding to the (3-5) neutron evaporation channels. Thus, calculations of survival probabilities with these values of a seem to be microscopically justified.