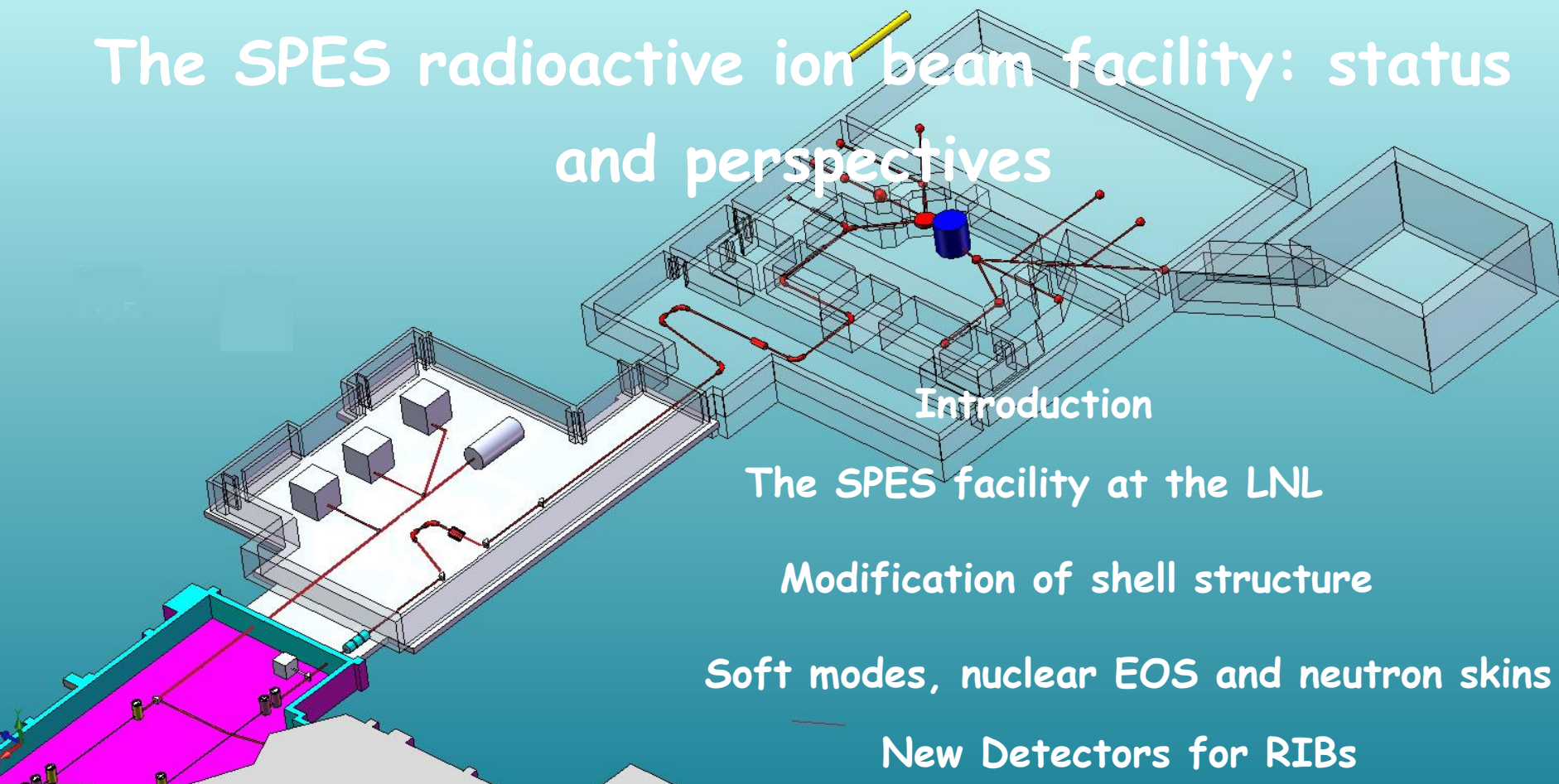


The SPES radioactive ion beam facility: status and perspectives



Introduction

The SPES facility at the LNL

Modification of shell structure

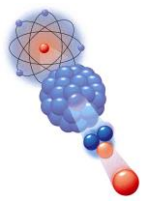
Soft modes, nuclear EOS and neutron skins

New Detectors for RIBs

Giacomo de Angelis

INFN Laboratori Nazionali di Legnaro

for The SPES collaboration



The goal of SPES

- **Selective Production of Exotic Species**

production of re-accelerated neutron-rich exotic beams
 10^{13} fission/s in-target production, and re-acceleration at $10 \cdot A$ MeV ($A=132$)

Radioisotope production & Medical applications
innovative radiopharmaceuticals (e.g. Sr-82, Cu-64, Cu-67)

- Optimized use of the two exits high current proton driver

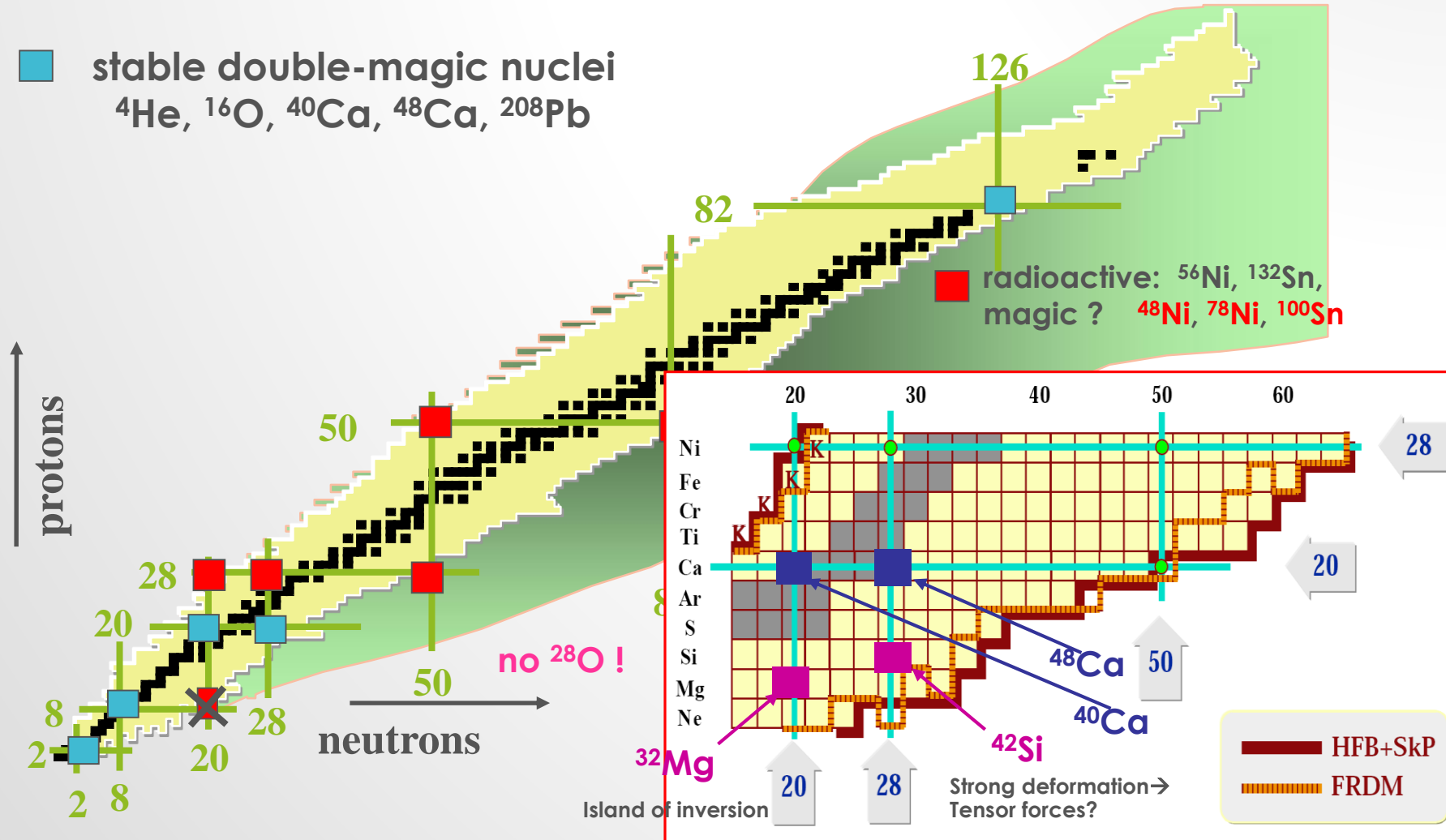
Fast neutron production & material applications: Atmospheric neutron spectra, QMN
Single Event Effect, neutron capture cross sections

SPES CHALLENGES

- Development of a comprehensive model of atomic nuclei – Do we understand the structure and stability of the nuclear systems?
- Understanding the origin of the elements and modeling of the extreme astrophysics environments
- Test of fundamental symmetries
- New applications of isotope science

One of the challenges: New magic numbers?

■ stable double-magic nuclei
 ${}^4\text{He}, {}^{16}\text{O}, {}^{40}\text{Ca}, {}^{48}\text{Ca}, {}^{208}\text{Pb}$



nature

THE INTERNATIONAL WEEKLY JOURNAL OF SCIENCE

Neutron number
34 makes exotic
calcium-54 isotopes
doubly magic

PAGE 207

MAGIC MOMENTS

NEUROSCIENCE

HERE'S LOOKING
AT MICE

Researchers at odds over
relevance of vision model

PAGE 156

CLIMATE

UNCHARTED
TERRITORY

When will global warming
top historical highs?

PAGES 174 & 183

EVOLUTION

COMING TO
A HEAD

The earliest
recognizable face?

PAGES 175 & 188

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10 October 2013 £10

Vol. 502, No. 7470



LETTER

doi:10.1038/nature12522

Evidence for a new nuclear 'magic number' from the level structure of ^{54}Ca

D. Steppenbeck¹, S. Takeuchi², N. Aoi³, P. Doornenbal², M. Matsushita¹, H. Wang², H. Baba², N. Fukuda², S. Go¹, M. Honma⁴, J. Lee⁵, K. Matsui⁶, S. Michimasa¹, T. Motobayashi⁷, D. Nishimura², T. Otsuka^{1,5}, H. Sakurai^{1,5}, Y. Shiga¹, P.-A. Söderström⁸, T. Sumikama⁹, H. Suzuki², R. Taniuchi⁵, Y. Utsuno⁹, J. J. Valiente-Dobón¹⁰ & K. Yoneda²

Atomic nuclei are finite quantum systems composed of two distinct types of fermion—protons and neutrons. In a manner similar to that of electrons orbiting in an atom, protons and neutrons in a nucleus form shell structures. In the case of stable, naturally occurring nuclei, large energy gaps exist between shells that fill completely when the proton or neutron number is equal to 2, 8, 20, 28, 50, 82 or 126 (ref. 1). Away from stability, however, these so-called 'magic numbers' are known to evolve in systems with a large imbalance of protons and neutrons. Although some of the standard shell closures can disappear, new ones are known to appear^{2,3}. Studies aiming to identify and understand such behaviour are of major importance in the field of experimental and theoretical nuclear physics. Here we report a spectroscopic study of the neutron-rich nucleus ^{54}Ca (a bound system composed of 20 protons and 34 neutrons) using proton knockout reactions involving fast radioactive projectiles. The results highlight the doubly magic nature of ^{54}Ca and provide direct experimental evidence for the onset of a sizeable subshell closure at neutron number 34 in isotopes far from stability.

The shell structure of the atomic nucleus was first successfully described more than 60 years ago¹. However, the question of how robust the standard magic numbers are in unstable nuclei with a large excess of neutrons—often referred to as 'exotic' nuclei—has been one of the main driving forces behind recent nuclear structure studies that focus on changes in the shell structure, called 'shell evolution'. A noteworthy example is the disappearance of the $N=28$ (neutron number 28) standard magic number in ^{48}Si (ref. 4), a nucleus that lies far from the stable isotopes on the Segrè chart. On the contrary, exotic oxygen isotopes⁵ provide evidence for the onset of a new shell closure at $N=16$, one that is not observed in stable nuclei. In both cases, the tensor force, a non-central component of the nuclear force, has a key role in describing the experimental spectra⁶.

The region of the Segrè chart around exotic calcium isotopes has also contributed valuable input to the understanding of nuclear shell evolution over recent years owing to experimental advances. Enhanced excitation energies of first $J^\pi = 2^+$ states (spin, J , parity, I) and reduced γ -ray transition probabilities, which are good indicators of nuclear shell gaps, for ^{52}Ca (refs 6, 7), ^{54}Ti (refs 8, 9) and ^{56}Cr (refs 10, 11) provide substantial evidence for the onset of a sizeable energy gap at $N=32$. This result was recently confirmed by high-precision mass measurements on neutron-rich Ca isotopes¹². In the framework of tensor-force-driven shell evolution⁶, the $N=32$ subshell closure is a direct consequence of the weakening of the attractive nucleon–nucleon interaction between protons (π) and neutrons (ν) in the $\pi f_{7/2}$ and $\nu f_{5/2}$ single-particle orbitals (SPOs) as the number of protons in the $\pi f_{7/2}$ SPO is reduced and the magnitude of the $\pi f_{7/2}$ – $\nu f_{5/2}$ energy gap increases (Fig. 1a–c).

A question that has been asked frequently over recent years is whether or not the onset of another subshell gap occurs in exotic

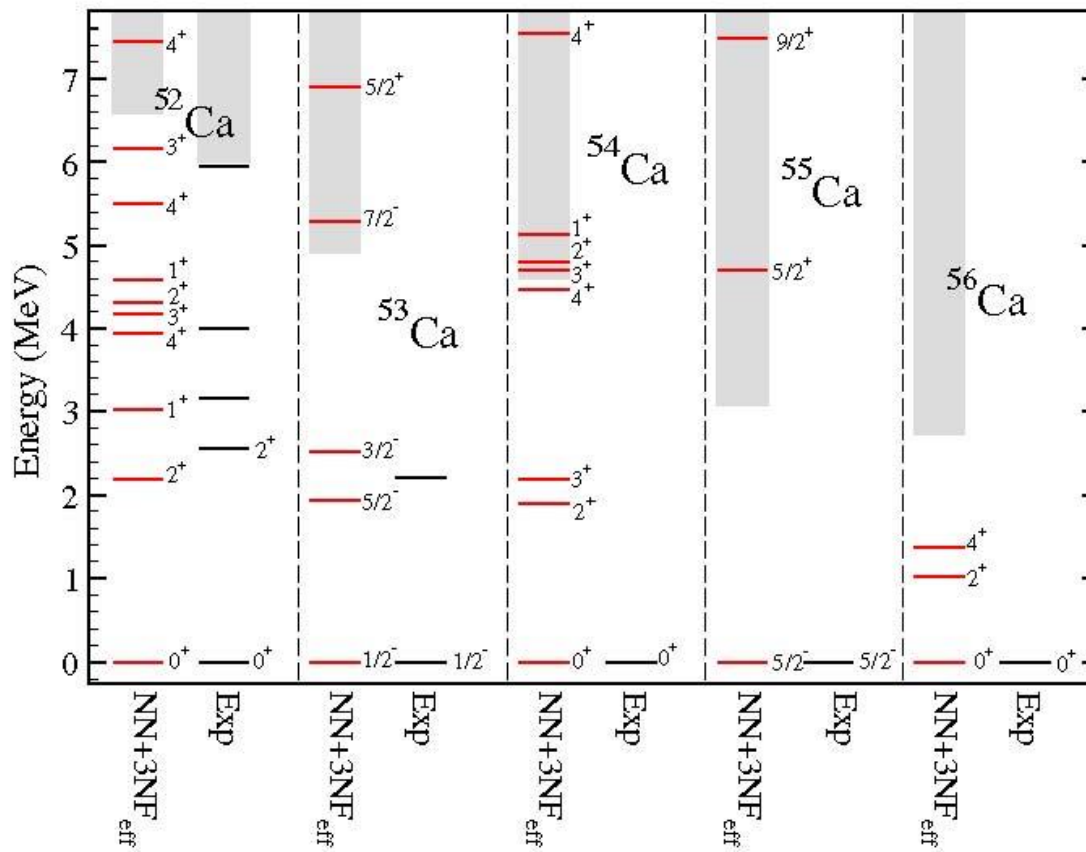
$N=34$ isotones, which was suggested qualitatively more than a decade ago¹³ on the basis of the general properties of nuclear forces. The onset of an appreciable subshell closure at $N=34$ is illustrated in Fig. 1d, indicating an energy gap between the $\nu p_{1/2}$ and $\nu f_{5/2}$ SPOs in ^{54}Ca that is comparable to the separation of the $\nu p_{3/2}$ and $\nu p_{1/2}$ spin-orbit partners, which is also implied by recent theoretical results; see, for example, ref. 14. We stress, however, that no $N=34$ subshell closure was reported in the experimental investigations of ^{56}Ti (refs 9, 15) or ^{58}Cr (refs 11, 16), and notable doubt on this magic number for Ca isotopes has been raised^{17,18}. Indeed, as indicated in Fig. 2a, theoretical predictions of the energy of the first $J^\pi = 2^+$ state for ^{54}Ca vary considerably, ranging from ~ 1 MeV in some cases to as high as ~ 4 MeV in others^{14,16,19–24}, despite exhibiting close agreement for lighter isotopes; for example, the predictions of the same theories lie within only 0.4 MeV of the empirical result for ^{52}Ca . Such stark discrepancies at $N=34$ reflect the need for direct experimental input on the matter.

To address this issue, we report on an experimental study of ^{54}Ca to clarify the strength of the $N=34$ subshell gap in nuclei farther from stability. The energies of nuclear excited states were investigated using proton knockout reactions involving ^{55}Sc and ^{56}Ti projectiles on a Be target at the Radioactive Isotope Beam Factory, Japan, operated by the RIKEN Nishina Center and the Center for Nuclear Study, University of Tokyo. Experimental details are provided in Methods Summary. Particle identification plots indicating the radioactive species transported through the BigRIPS separator and ZeroDegree spectrometer²⁵, which were used to select and tag radioactive beam projectiles and reaction products, are presented in Fig. 3a and Fig. 3b, respectively. We emphasize that the intensity of the radioactive beam reported here, which was critical to the success of the experiment, is unique to the Radioactive Isotope Beam Factory. Excited-state energies were deduced using the technique of in-beam γ -ray spectroscopy.

The γ -rays measured in coincidence with ^{54}Ca projectiles produced through the one- and two-proton knockout reaction channels are presented in Fig. 4a. The γ -ray energies measured in the laboratory frame of reference have been corrected for Doppler shifts, and so the transitions appear at the energies they would be in the rest frame of the nucleus. The most intense γ -ray line in the ^{54}Ca spectrum, the peak at 2,043(19) keV (error, 1 s.d.) in Fig. 4a, is assigned as the transition from the first 2^+ state (2_1^+) to the 0^+ ground state. In addition, two weaker transitions are located at 1,656(20) and, respectively, 1,184(24) keV. Figure 4b shows a γ -ray spectrum obtained with the condition of a prompt coincidence (≤ 10 ns) with the 2,043-keV γ -ray, indicating that the weaker transitions were emitted in decay sequences involving the $2_1^+ \rightarrow 0^+$ ground-state transition. On the basis of the γ -ray relative intensities, the 1,656-keV transition is proposed to depopulate a level at 3,699(28) keV, as presented in the ^{54}Ca level scheme in the lower-right section of Fig. 4a. Placement of the 1,184-keV transition in the

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N=34 subshell closure due to the effects of three body forces driving the monopole part of the nuclear Hamiltonian



Beta delayed
n spectroscopy
n- γ coincidences

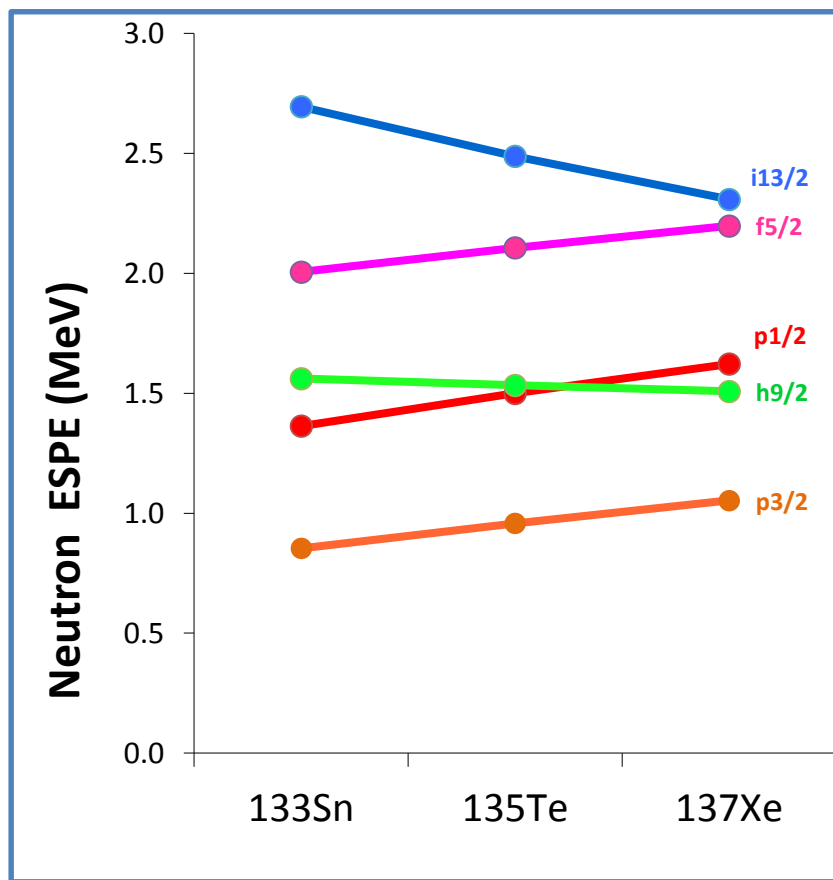
LOI SPES A. Gottardo IPN Orsay

Shell model calculations with effective interaction based on chiral Effective field theory and three body forces (G. Hagen PRL 109 2012)

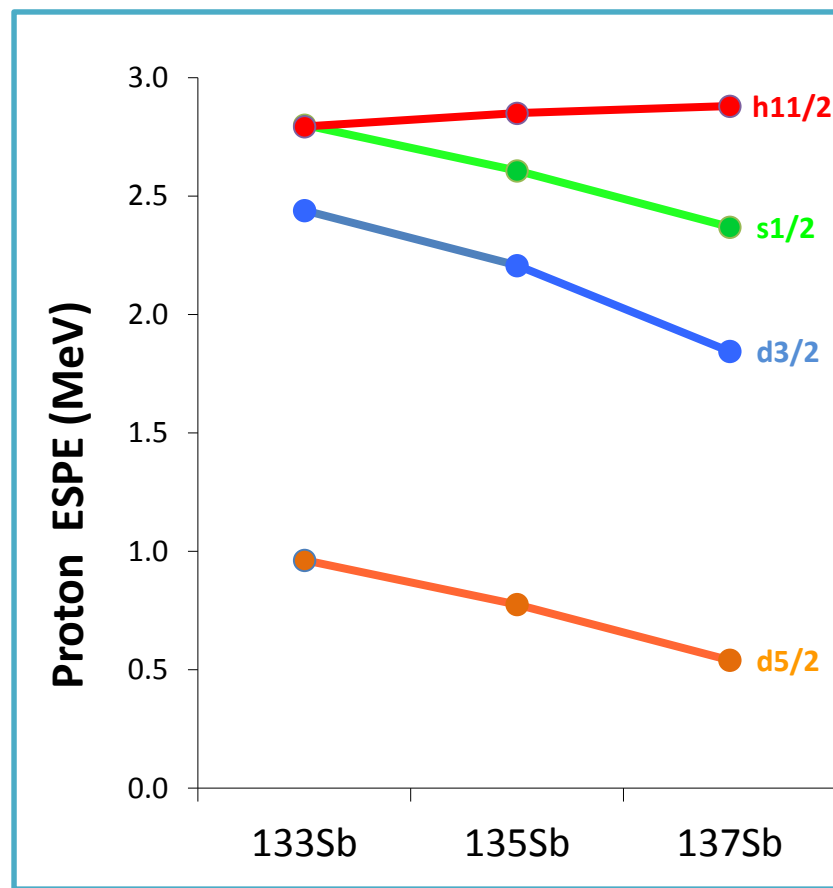
Evolution of the single-particle states around ^{132}Sn

$$\bar{\varepsilon}_{j_v} = \varepsilon_{j_v} + \sum_{j_\pi} V^M(j_v j_\pi) N_{j_\pi}$$

$$\bar{\varepsilon}_{j_\pi} = \varepsilon_{j_\pi} + \sum_{j_v} V^M(j_\pi j_v) N_{j_v}$$



Courtesy of A. Gargano

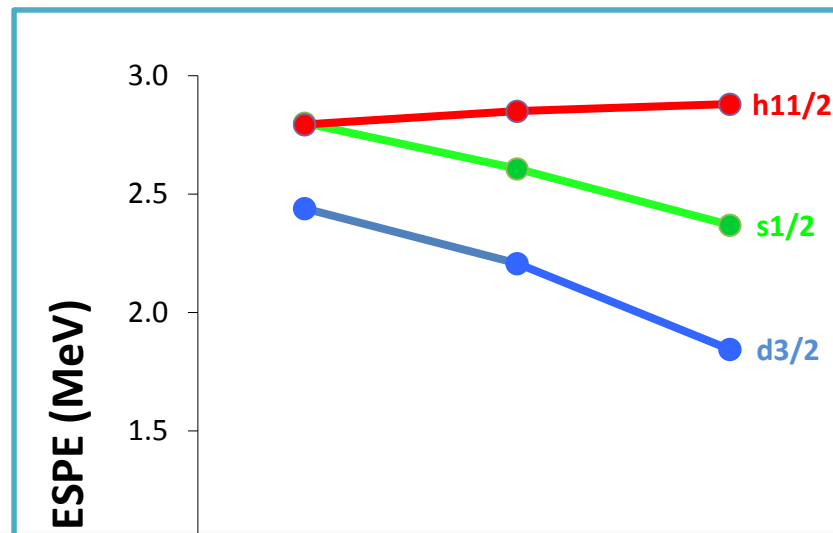
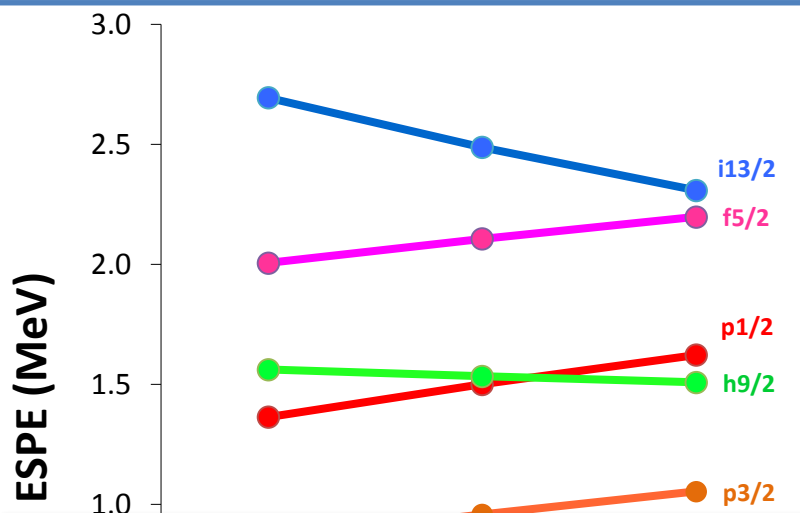


LOI SPES D. Mengoni (Uni. Pd)

Evolution of the single-particle states around ^{132}Sn

$$\bar{\varepsilon}_{j_v} = \varepsilon_{j_v} + \sum_{j_\pi} V^M(j_v j_\pi) N_{j_\pi}$$

$$\bar{\varepsilon}_{j_\pi} = \varepsilon_{j_\pi} + \sum_{j_v} V^M(j_\pi j_v) N_{j_v}$$



One-particle spectroscopic factors which give direct information on single-particle excitations

0.0 ^{133}Sn ^{135}Te ^{137}Xe

0.0 ^{133}Sb ^{135}Sb ^{137}Sb

Courtesy of A. Gargano

LOI SPES D. Mengoni (Uni. Pd)

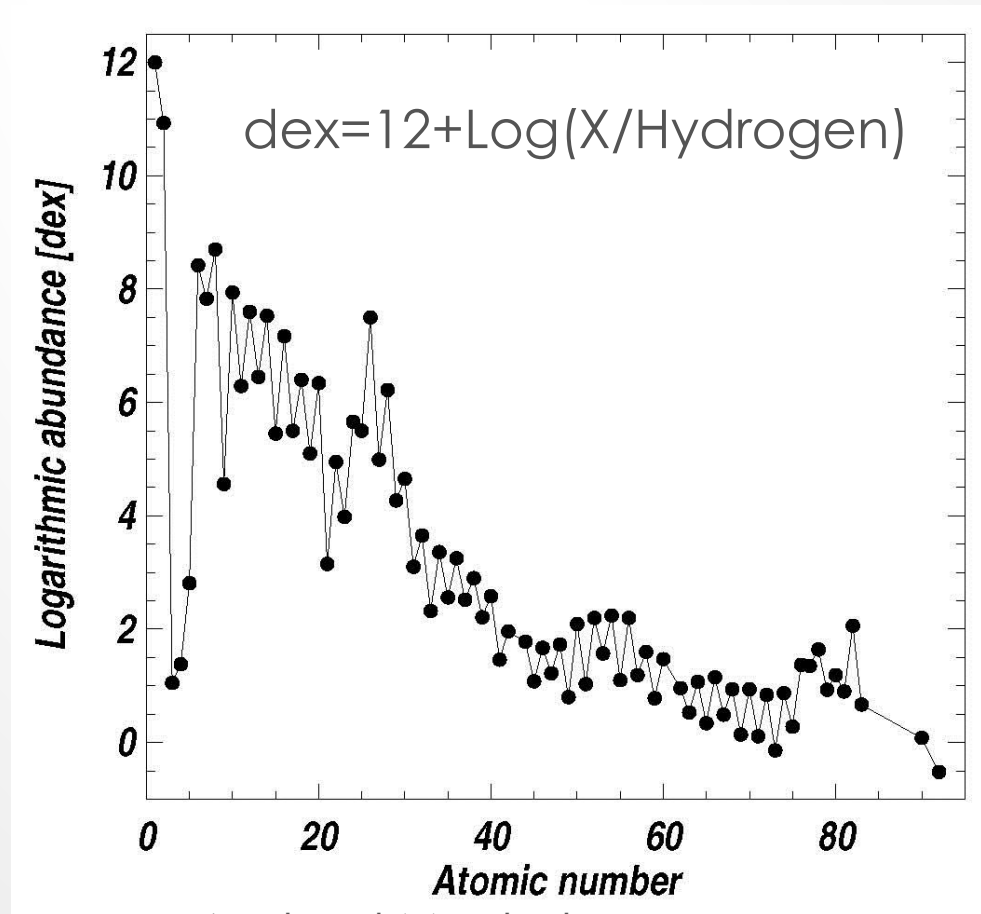
ONE OF THE CHALLENGES: ORIGIN OF THE ELEMENTAL ABUNDANCES IN THE SOLAR SYSTEM

Stars are mostly made of hydrogen and helium, but each has a fairly unique pattern of other elements

The abundance of elements tells us about the history of events prior to the formation of our Sun

The plot shows the composition of the Sun Photosphere

How are these elements created prior to the formation of the Sun?



Asplund M. et al.
Ann. Rev. Astr. 47, 481 (2009)

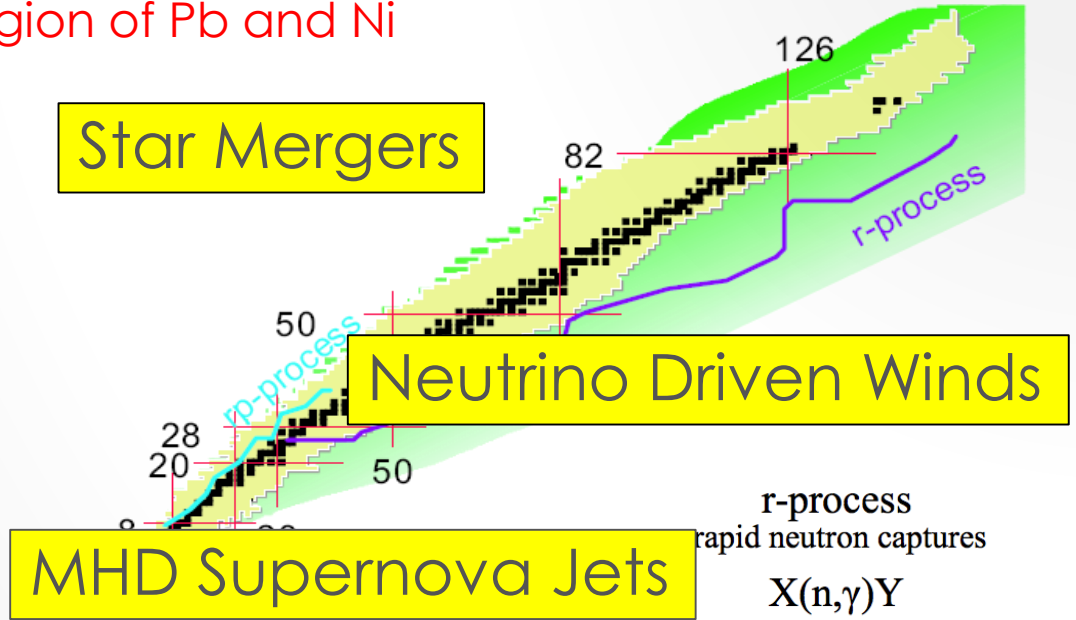
n-rich region of Pb and Ni

- β-decay half-lives
- neutron emission probabilities

different in **DEFORMED** and **SPHERICAL** nuclei

$$\tau_{\beta \text{ SPH}} \sim 10 \times \tau_{\beta \text{ DEF}}$$

$$P_{n \text{ SPH}} \sim 0.5 \times P_{n \text{ DEF}}$$



- Large uncertainty in r-process location
- difficulty to extrapolate to more exotic regions

Importance of Nuclear Structure studies in the vicinity/towards r-process path:

- ... first excited states
- ... Isomeric states
- ... β decay lifetimes

ONE OF THE CHALLENGES: REFLECTION ASYMMETRIC NUCLEI AND STATIC ELECTRIC DIPOLE MOMENT



The lopsided nuclei, described today (May 8) in the journal Nature, could be good candidates for researchers looking for new types of physics beyond the reigning explanation for the bits of matter that make up the universe (called the [Standard Model](#)), said study author [Peter Butler](#), a physicist at the University of Liverpool in the United Kingdom.

The findings could help scientists search for physics beyond the Standard model, said [Witold Nazarewicz](#). An electric dipole moment would provide a way to test extension theories to the Standard Model, such as supersymmetry, which could help explain why there is more matter than antimatter in the universe.

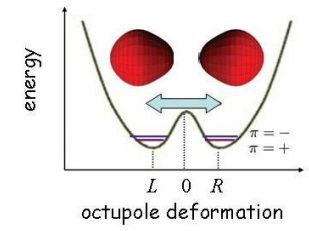
ONE OF THE CHALLENGES: REFLECTION ASYMMETRIC NUCLEI AND STATIC ELECTRIC DIPOLE MOMENT

V Spevak, N Auerbach, and VV Flambaum
PR C 56 (1997) 1357

Schiff moment:

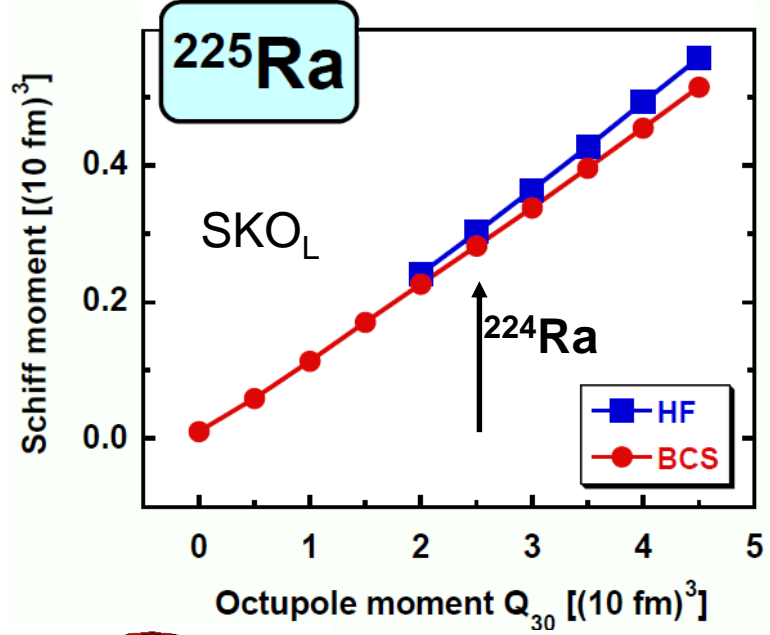
$$S = -2 \frac{J}{J+1} \frac{\langle \hat{S}_z \rangle \langle \hat{V}_{PT} \rangle}{\Delta E}$$

related to Q_3 P,T-violating n-n interaction

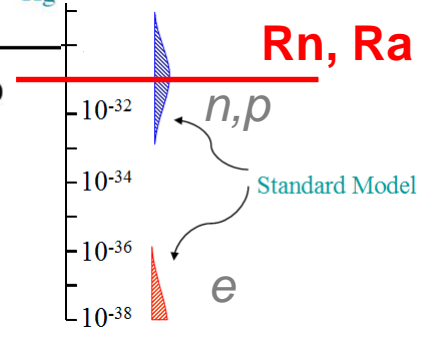
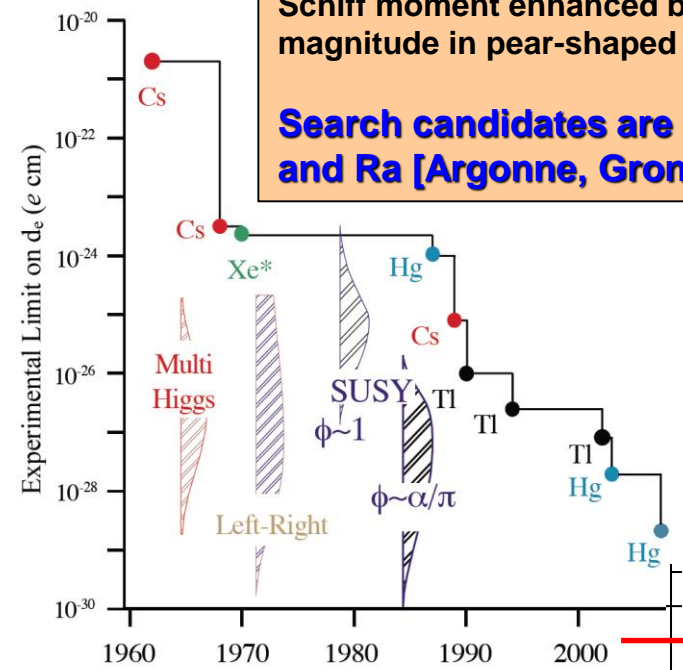


energy splitting of parity doublet

J Dobaczewski (Trento, 2010)



Schiff moment enhanced by ~ 3 orders of magnitude in pear-shaped nuclei
Search candidates are odd-A Rn [TRIUMF] and Ra [Argonne, Groningen/ISOLDE]

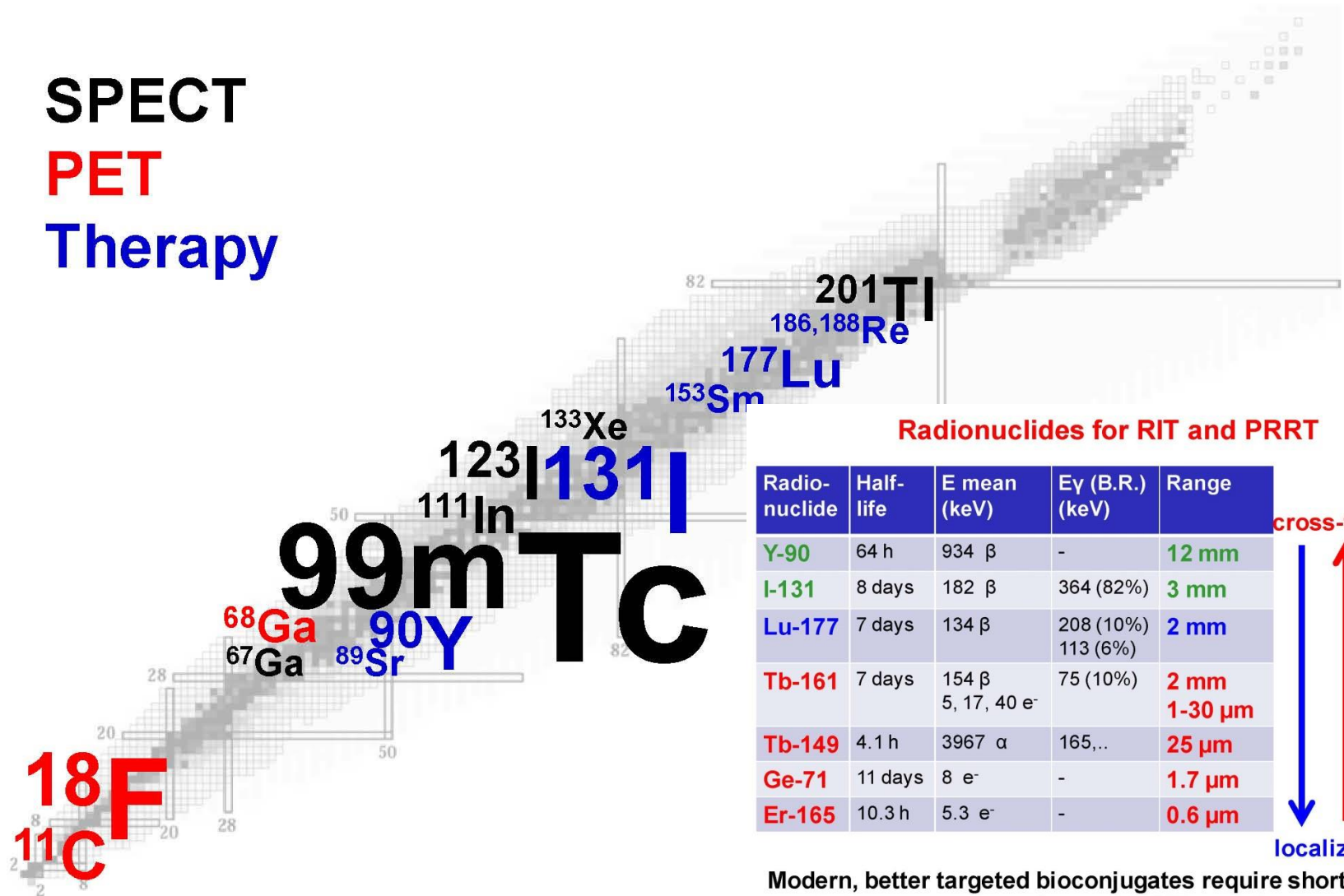


Measure: Q_3 in even-A Rn, Ra
 ΔE in odd-A Rn

ONE OF THE CHALLENGES: RADIONUCLIDES FOR MEDICINE

The chart of nuclides – nuclear medicine perspective

SPECT
PET
Therapy



Radionuclides for RIT and PRRT

Radio-nuclide	Half-life	E mean (keV)	E _γ (B.R.) (keV)	Range
Y-90	64 h	934 β	-	12 mm
I-131	8 days	182 β	364 (82%)	3 mm
Lu-177	7 days	134 β	208 (10%) 113 (6%)	2 mm
Tb-161	7 days	154 β 5, 17, 40 e ⁻	75 (10%)	2 mm 1-30 μm
Tb-149	4.1 h	3967 α	165,...	25 μm
Ge-71	11 days	8 e ⁻	-	1.7 μm
Er-165	10.3 h	5.3 e ⁻	-	0.6 μm

cross-fire
↑
Established isotopes
Emerging isotopes
↓
R&D isotopes: supply-limited!
localized

Modern, better targeted bioconjugates require shorter-range radiation ⇒ need for **adequate (R&D) radioisotope supply**.



Second SPES International Workshop

26-28 May 2014 *INFN Laboratori Nazionali di Legnaro*
Europe/Rome timezone

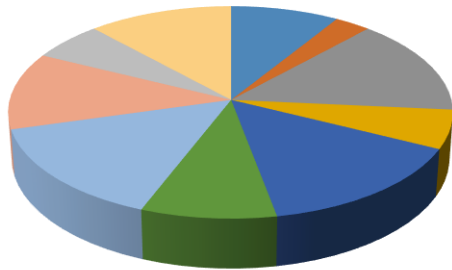
Presented 37 Letters of Intent

SPES2010 Workshop

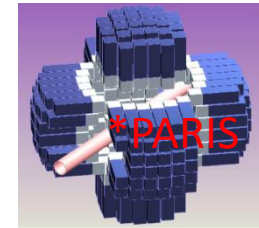
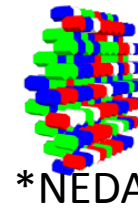
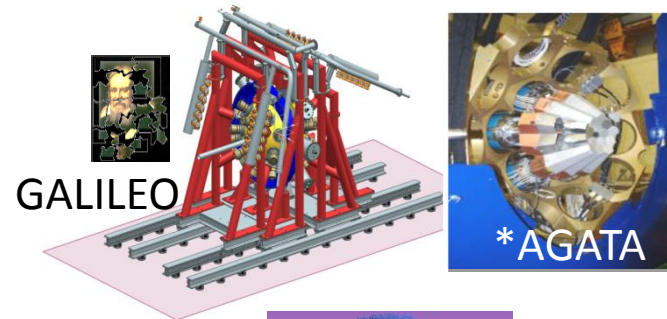
(LNL- November 15th-17th, 2010)

24 Lol's for reaccelerated exotic beams

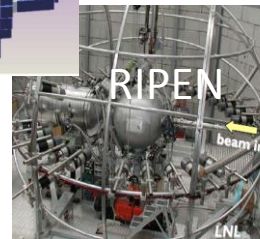
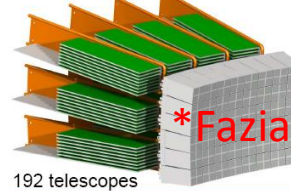
SPES LOIs Topics



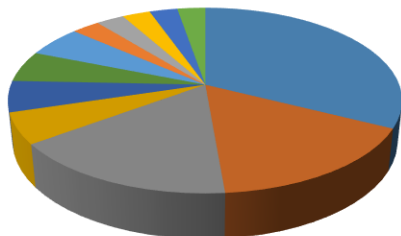
- GS properties
- moments
- Coulex
- DirReac with ActiveTarget
- DirReac with Si
- Mn transfer



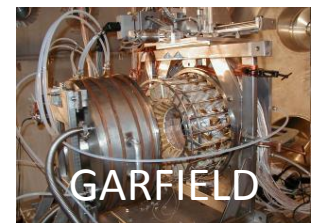
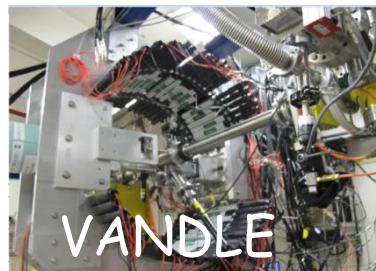
PHASE-II DEMONSTRATOR




SPES LOIs Spokespersons



- Italy
- France
- Poland
- Russia
- USA
- Belgium
- Croatia
- Norway
- Bulgaria
- Spain
- Russia
- China



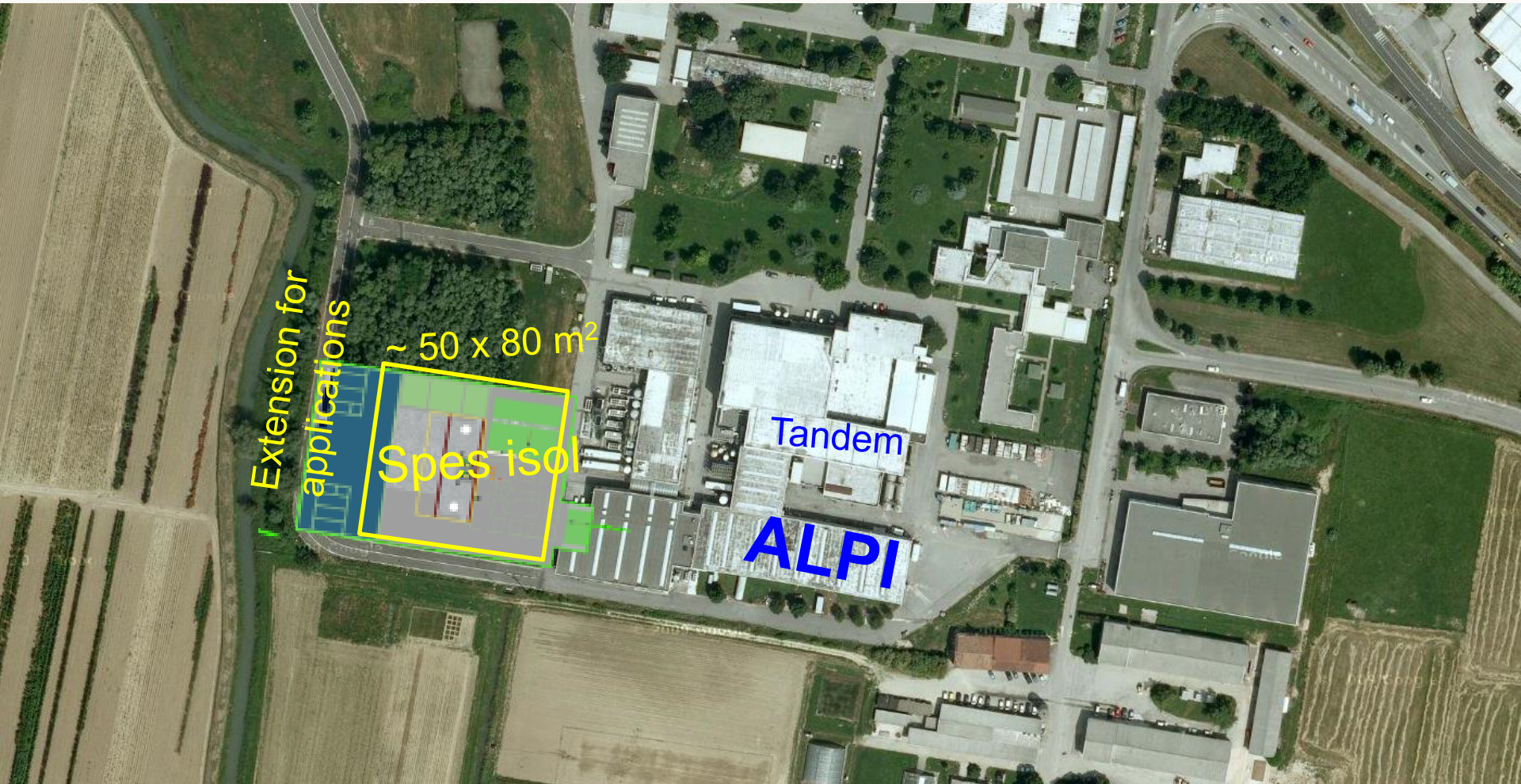
LOIs at ISOL Facilities

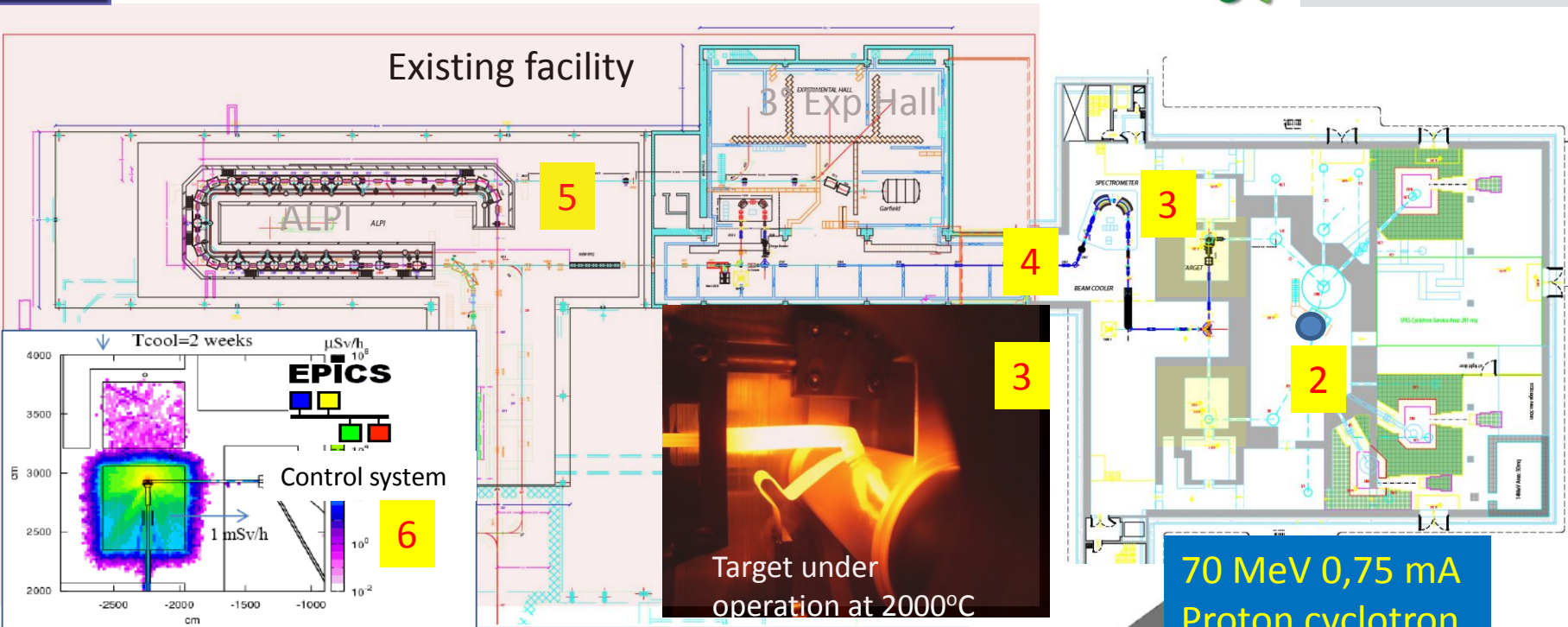


	SPES	SPIRAL2 Phase 2	HIE-ISOLDE
Decay Studies	4	20	Not Applicable
Elastic	2	4	3
COULEX	7	2	13
Transfer	8 (3HI)	7	16
Deep Inelastic/ MNTR	5	1	0
Fusion/Fission	11	3	2
New instrumentation	4	NA	5
Astrophysics	3	6	4



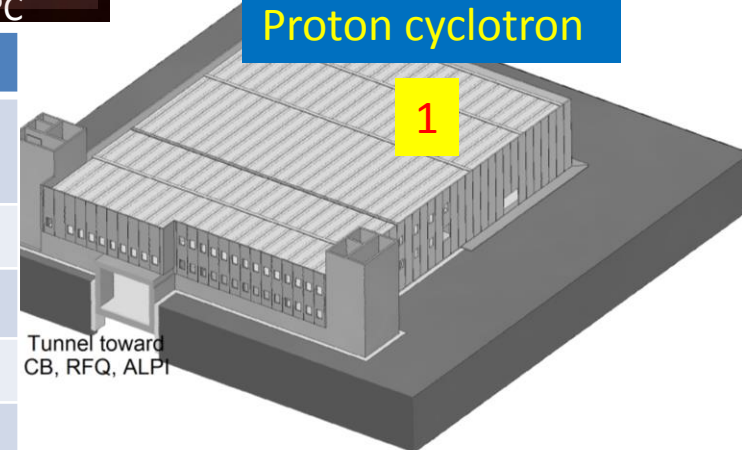
SPES Facility @ LNL

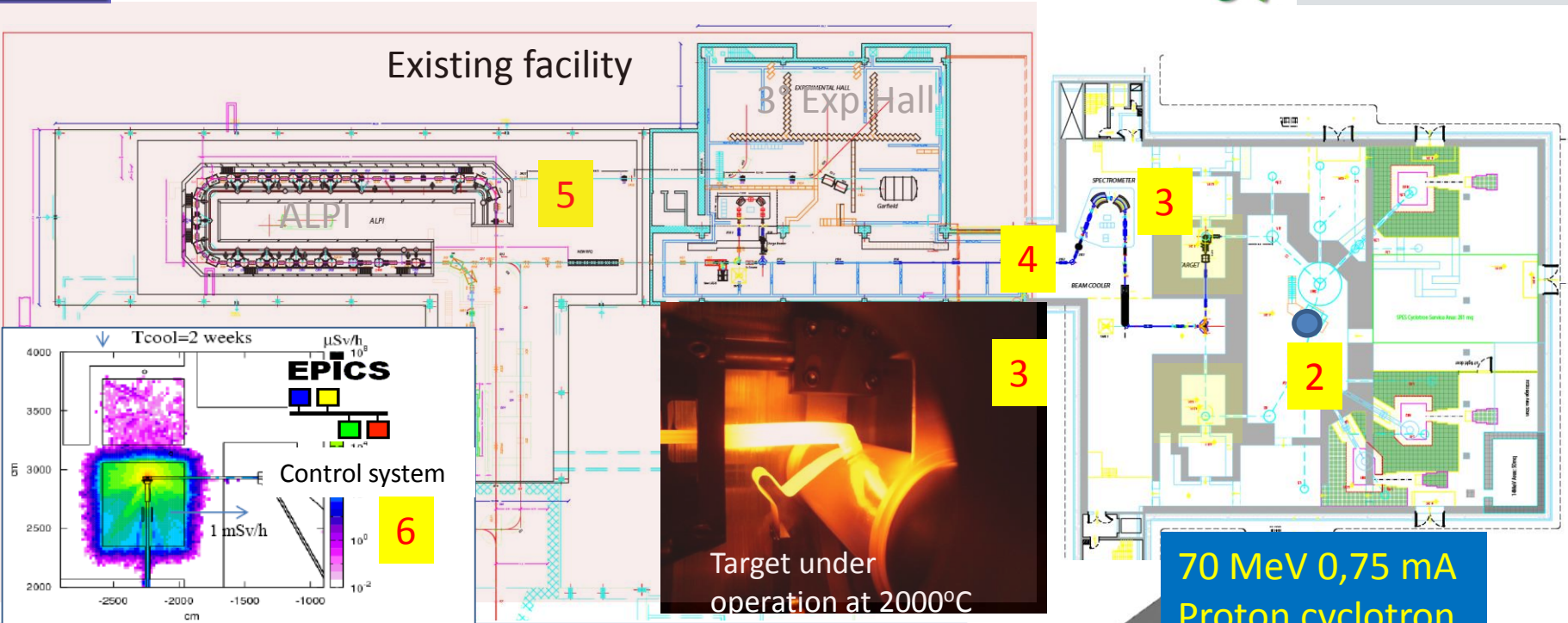




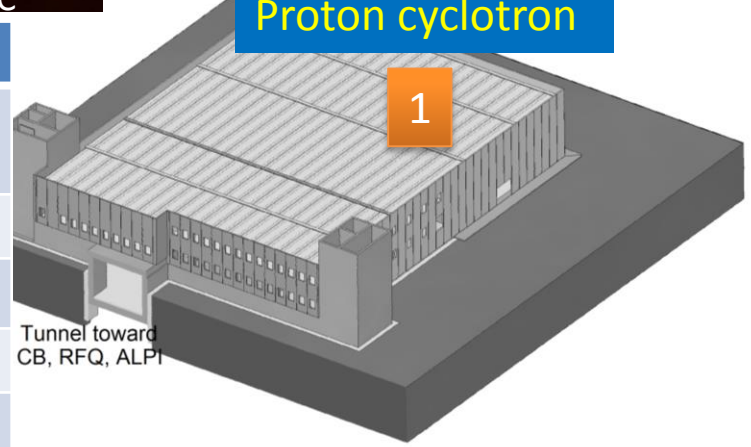
SPES sub-systems

- 1 Building and infrastructures with 2 ISOL bunkers for radioactive beam and application area for radioisotopes and neutrons
- 2 Cyclotron 70 MeV protons with 2 independent exits
- 3 ISOL UCx target designed for 10^{13} f/s
- 4 Beam transport with High Resolution Mass Separation
- 5 Reacceleration with ALPI superconducting linac (10A MeV A=130)
- 6 Radioprotection, safety & controls





70 MeV 0,75 mA
Proton cyclotron



Tunnel toward
CB, RFQ, ALPI

SPES sub-systems

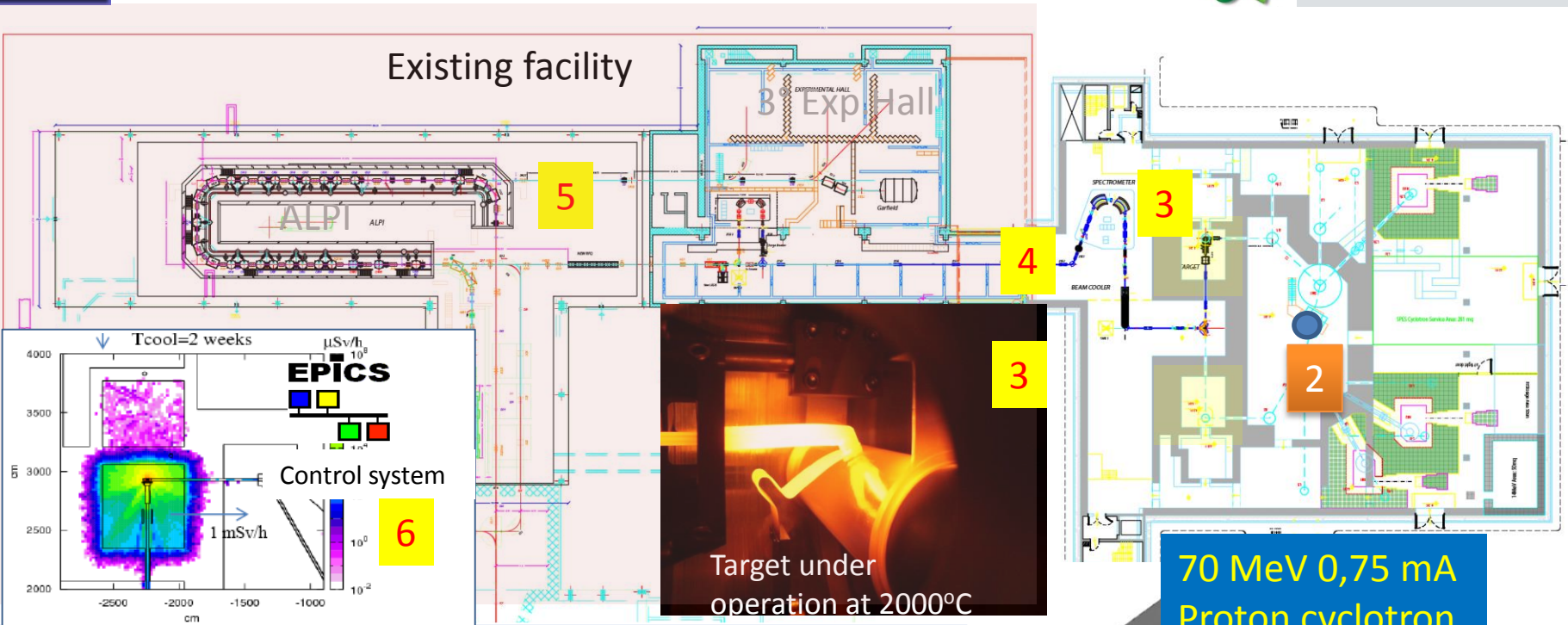
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<http://www.lnl.infn.it/~gestimp/speslive.jpg>

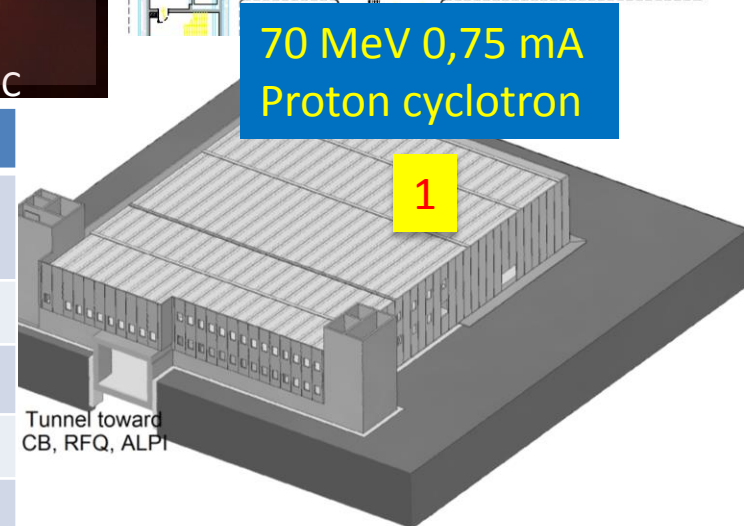


SPES Facility Layout



SPES sub-systems

- 1 Building and infrastructures with 2 ISOL bunkers for radioactive beam and application area for radioisotopes and neutrons
- 2 Cyclotron 70 MeV protons with 2 independent exits
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- 6 Radioprotection, safety & controls



70 MeV 0,75 mA
Proton cyclotron

Cyclotron test at BEST Company site (Ottawa)

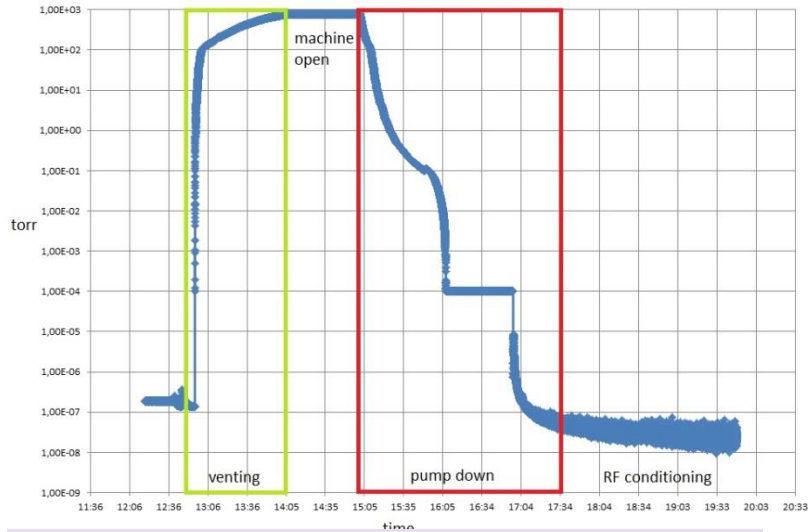


Main Parameters

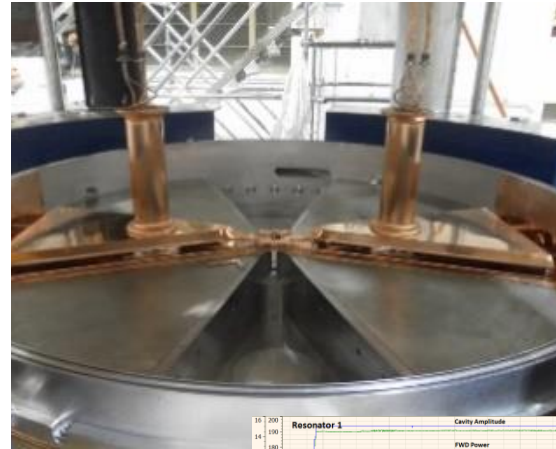
Accelerator Type	Cyclotron AVF 4 sectors
Particle	Protons (H⁺ accelerated)
Energy	Variable within 30-70 MeV
Max Current Accelerated	750 μA (52 kW max beam power)
Available Beams	2 beams at the same energy (upgrade to different energies)
Max Magnetic Field	1.6 Tesla
RF frequency	56 MHz, 4 th harmonic mode
Ion Source	Multicusp H ⁺ I=15 mA, Axial Injection
Dimensions	$\Phi=4.5$ m, h=1.5 m
Weight	150 tons

Cyclotron assembled and operated at 1MeV

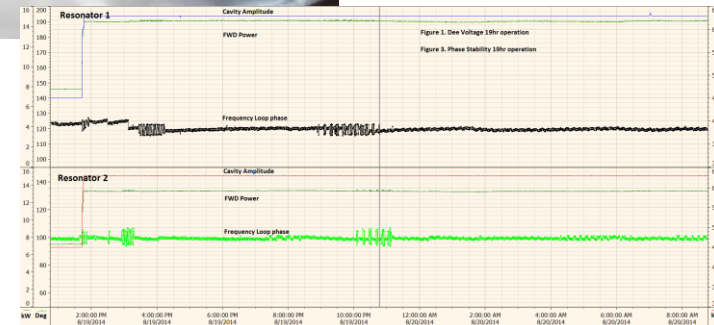
Final Factory Acceptance Test in Ottawa next week



Successful vacuum test: below 10^{-7} Torr
after 2 hours pumping

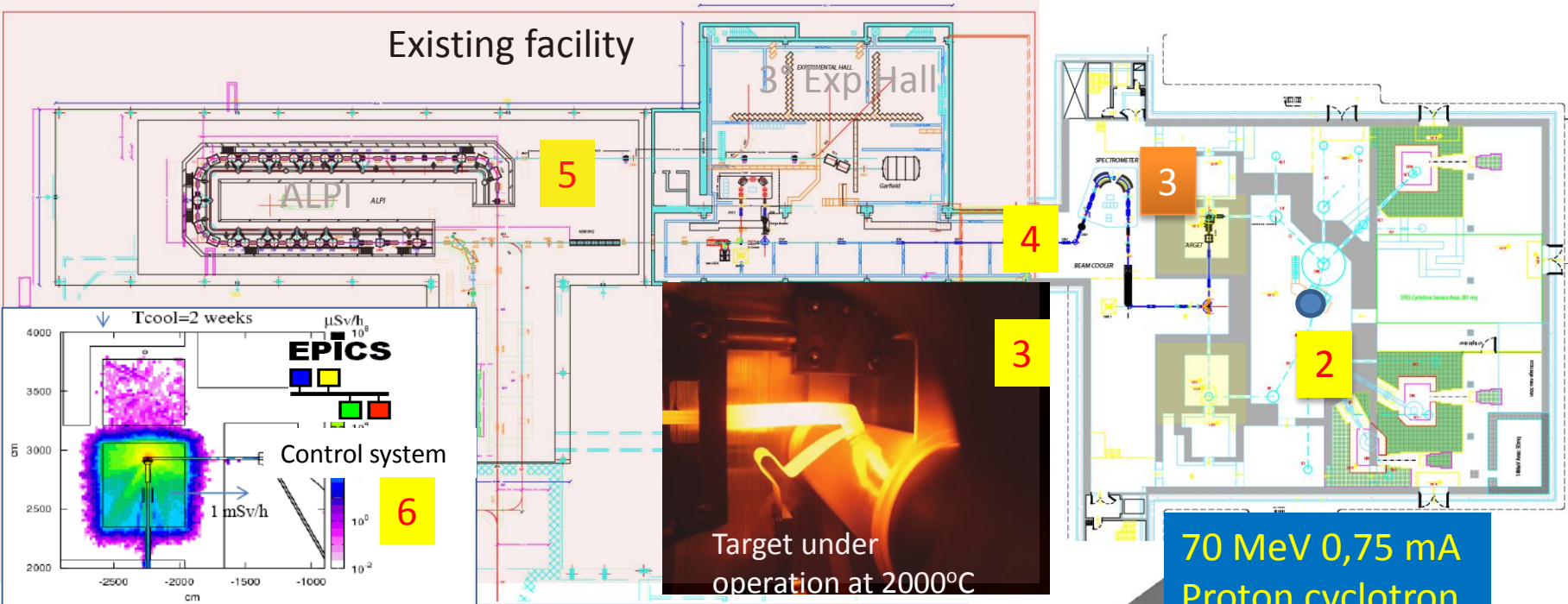


Both cavities with
62 kV voltage and
14 kW FWD RF
power each.
Amplitude Stability
within $\pm 2.5 \cdot 10^{-5}$

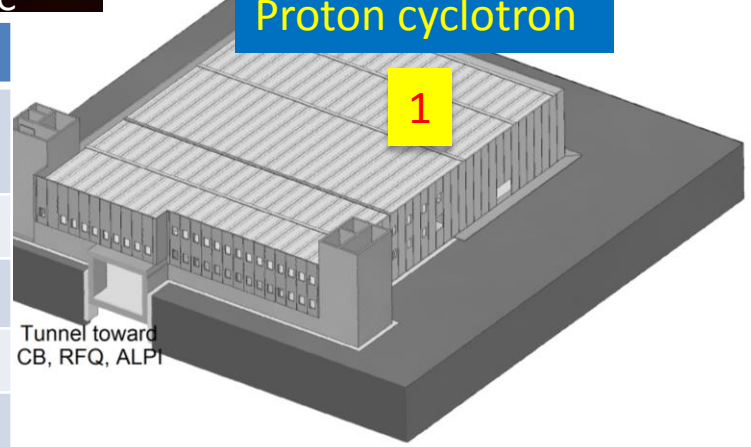


Critical point: maximum current of 800 microA accelerated up to 1 MeV
If test is successful, the cyclotron will be dismantled and the transfer to Italy will start.

Cyclotron at LNL: February 28, 2015



70 MeV 0,75 mA
Proton cyclotron



Tunnel toward
CB, RFQ, ALPI

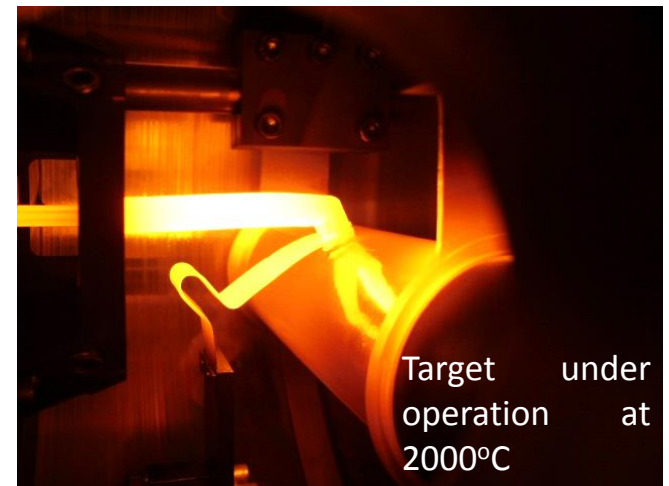
SPES sub-systems

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- 6 Radioprotection, safety & controls

SPES DIRECT TARGET CONCEPT to operate with **8 kW** proton beam

(A. Andrichetto et al.)

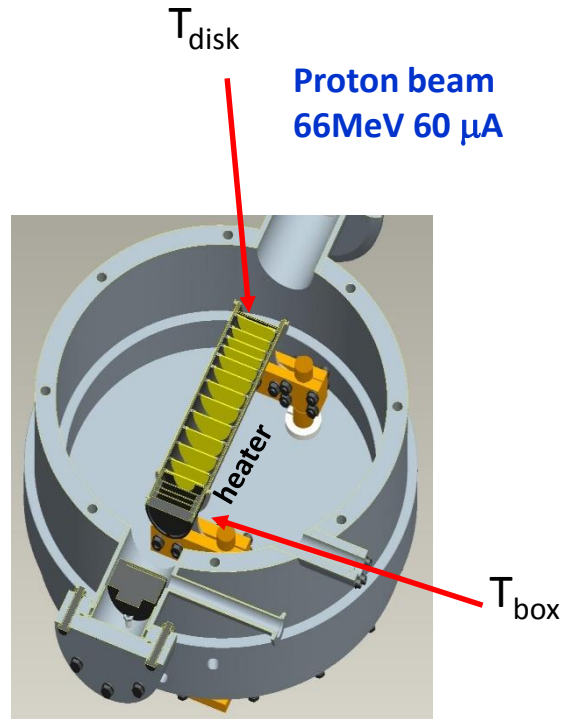
- Direct Target carefully designed to reach **10^{13} fissions/s** with **8 kW** proton beam (thermo-mechanical considerations);
- **In beam test** performed at **iThemba lab (South Africa)** on May 2014;
- Prototype under operation.
- Fully developed **front-end** following ISOLDE design;



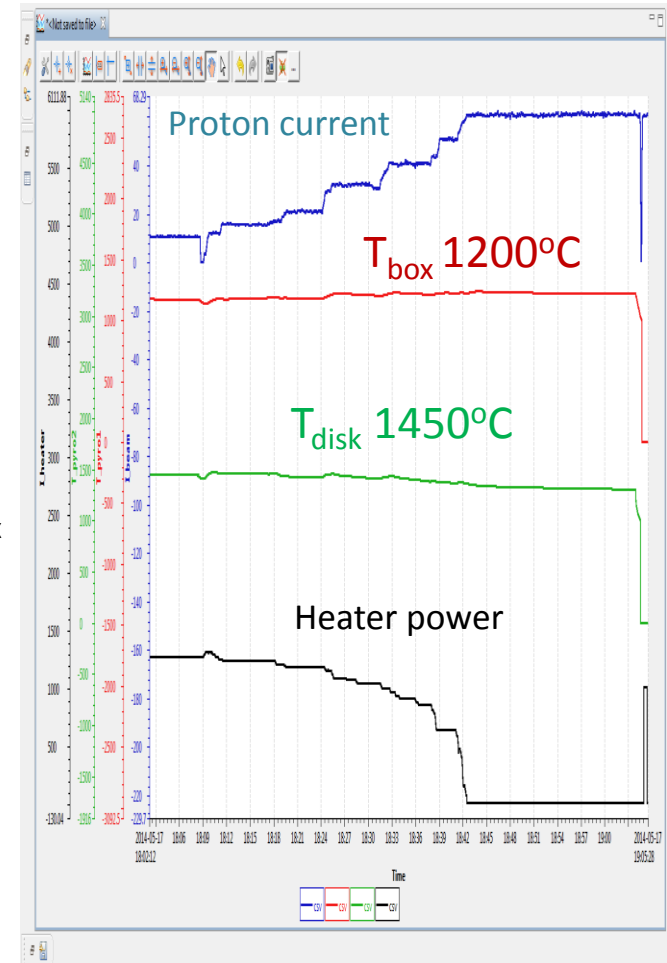
SPES target in-beam power test (SiC target)

Heater power compensated by proton beam.

- Up to **4 kW proton beam** in target.
- **Stable temperatures**
- **Stable vacuum** ($3 \cdot 10^{-5}$ mbar)

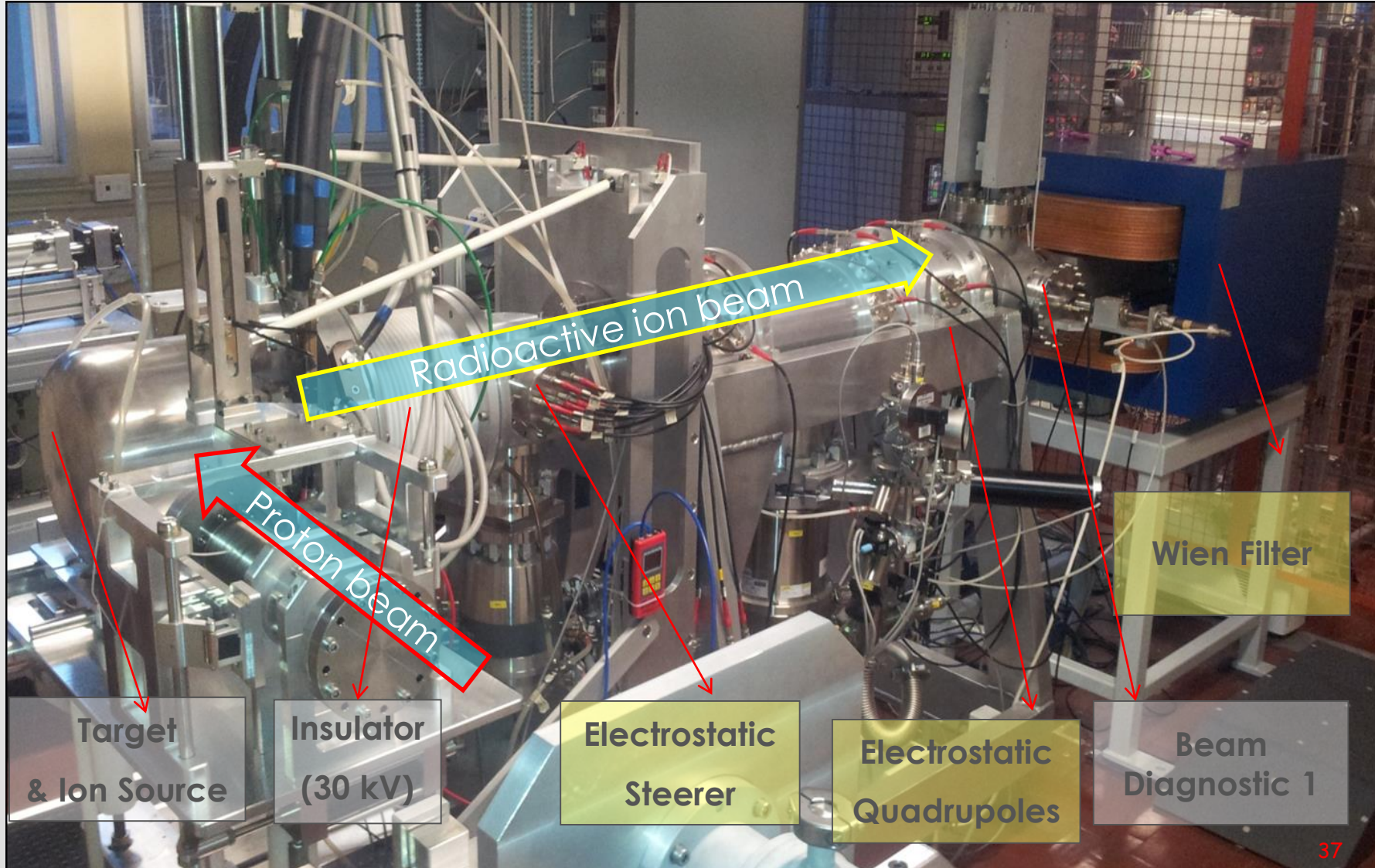


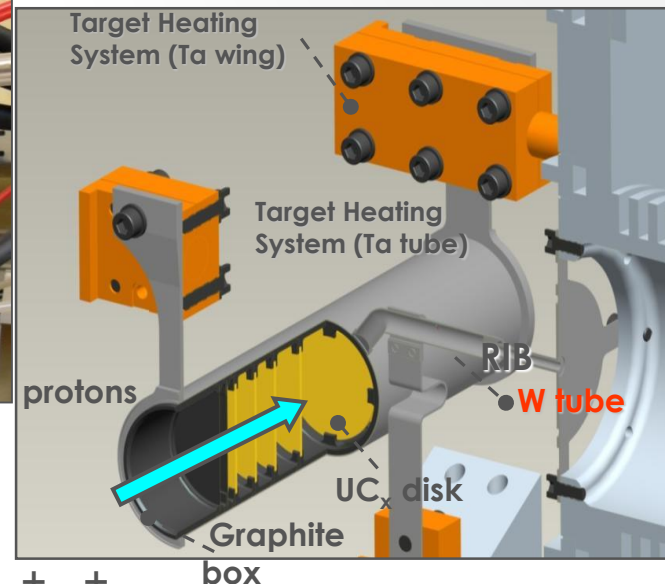
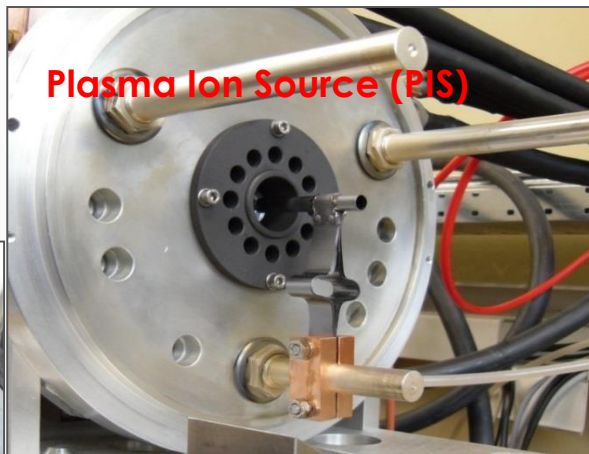
Thanks to Rob, Lowry and all the
iThemba_Labs Cyclotron staff



iThemba_Labs, May 17th, 2014

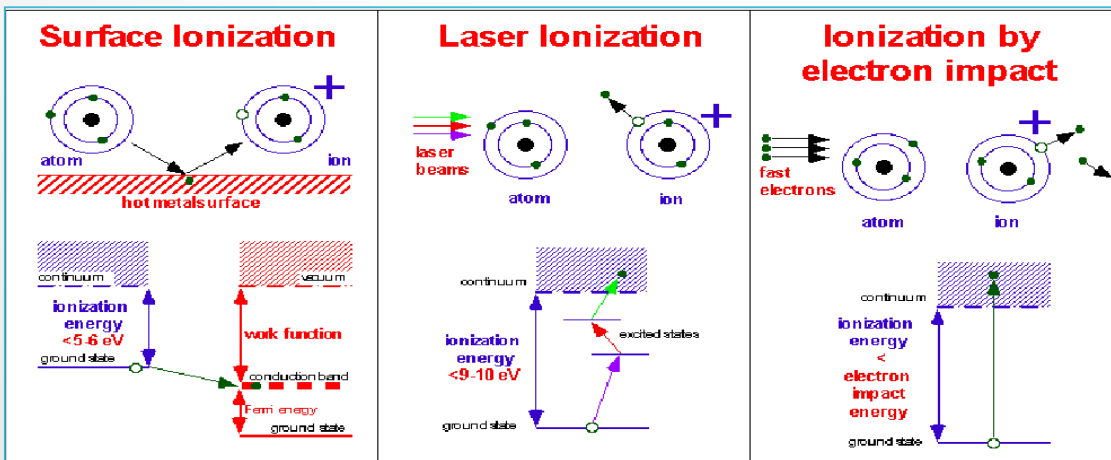
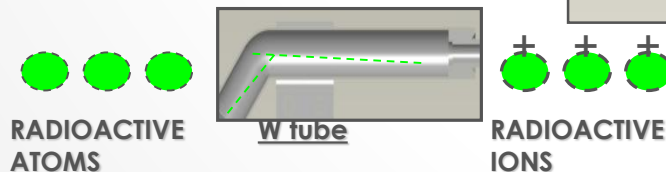
(working since 2010)

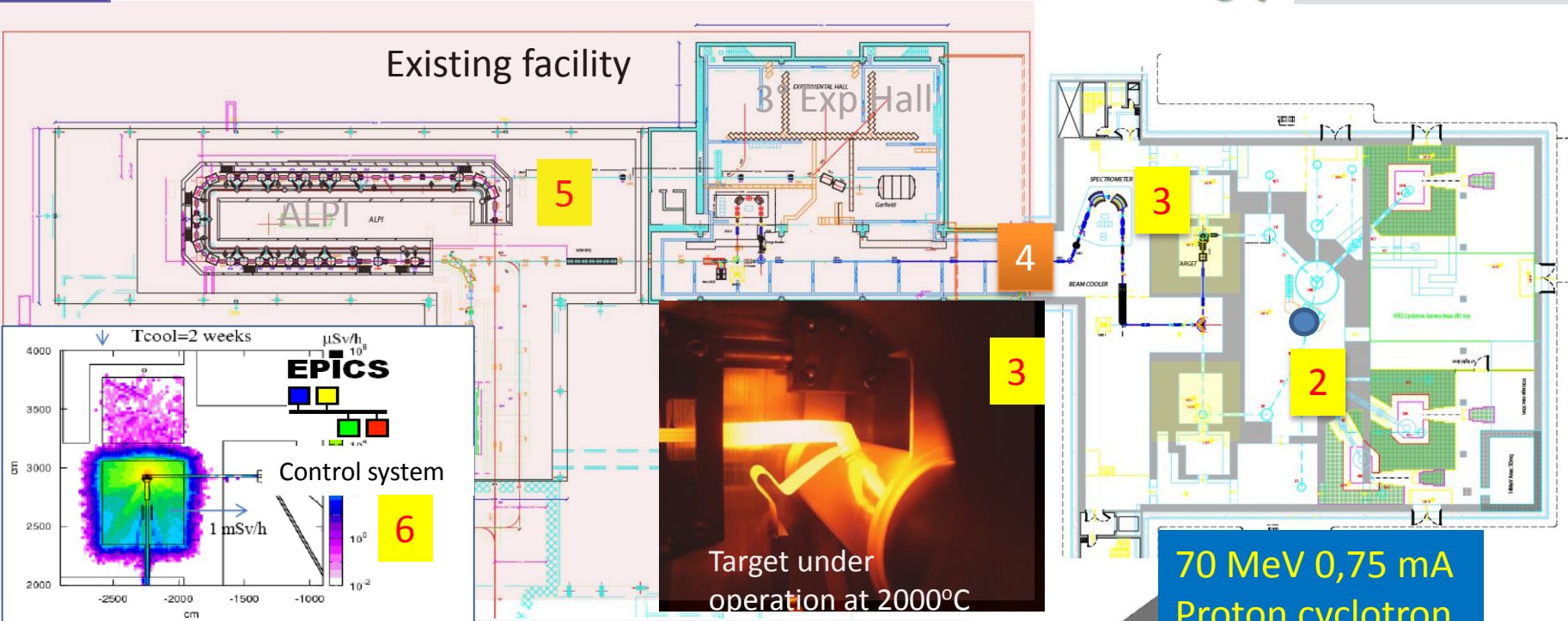




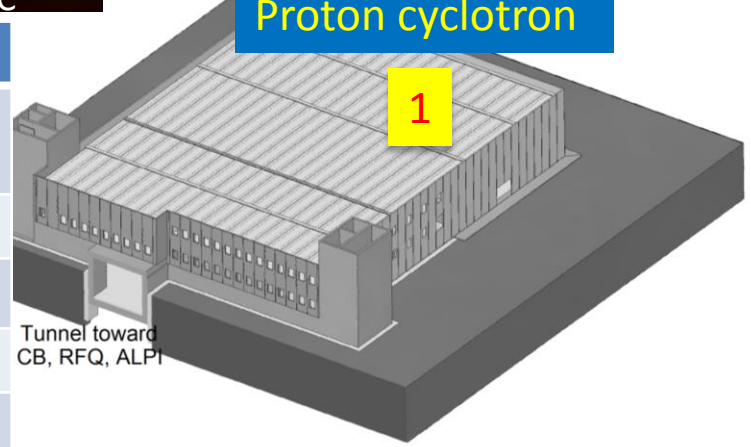
ION SOURCE DEVICE

Transfer Line





70 MeV 0,75 mA
Proton cyclotron



SPES sub-systems

- 1 Building and infrastructures with 2 ISOL bunkers for radioactive beam and application area for radioisotopes and neutrons
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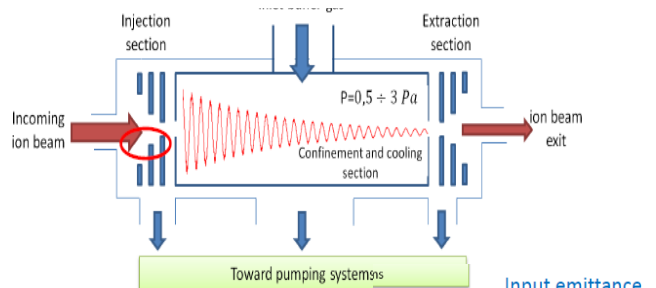
L.Calabretta, M.Comunian, **Collaboration SPES – CENBG Bordeaux (SPIRAL2)**
A.Russo, L.Bellan

- Scaled-up version of the separator designed by Cary Davids for CARIBU, Argonne
- Mass resolution: 1/40000 (eng. design: 1/25000)

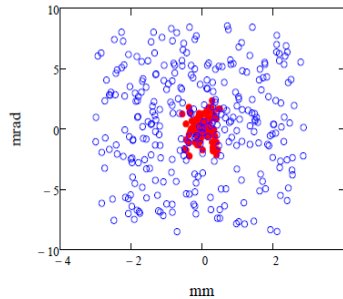
Beam Cooler to match the HRMS input requirements

COOLBEAM experiment financed by INFN-CSN5, 2012→2015
Collaboration: LNL-LNS, Mi bicocca

M.Maggiore



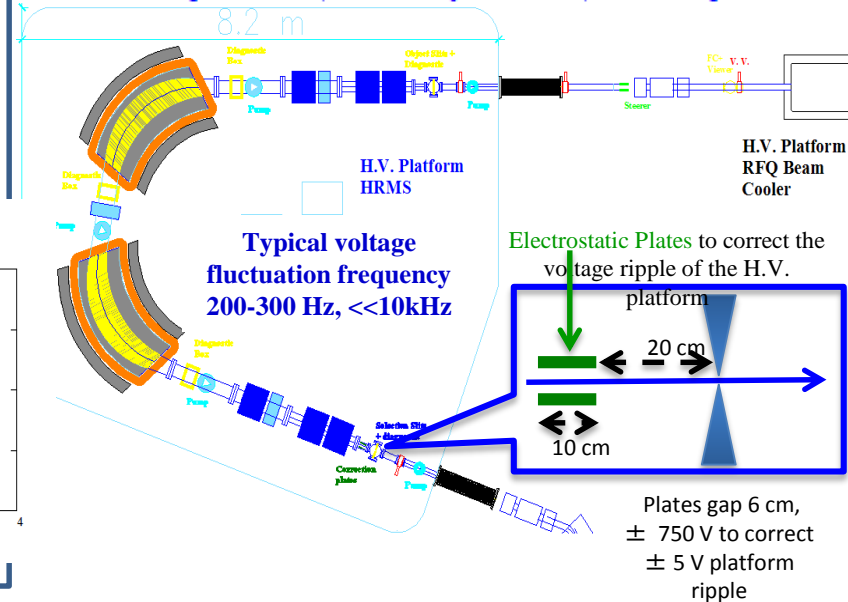
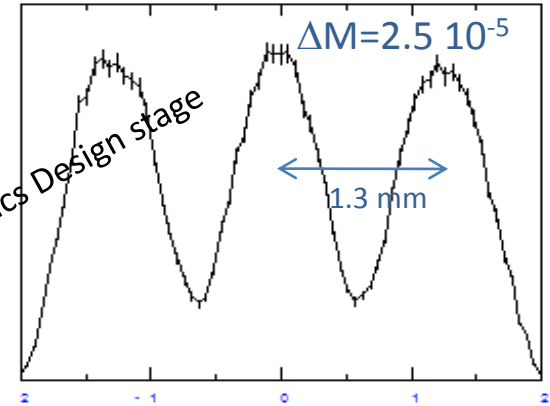
Input emittance
Output emittance



ELEMENTO	QTA	NUMERO PARTI	DESCRIZIONE
1	10	1	Struttura
2	10	1	Condensatore isolante
3	10	1	Struttura isolante
4	10	1	Struttura isolante
5	10	1	Struttura isolante
6	10	1	Struttura isolante
7	10	1	Struttura isolante
8	10	1	Struttura isolante
9	10	1	Struttura isolante
10	10	1	Struttura isolante
11	10	1	Struttura isolante
12	10	1	Struttura isolante
13	10	1	Struttura isolante
14	10	1	Struttura isolante
15	10	1	Struttura isolante
16	10	1	Struttura isolante
17	10	1	Struttura isolante
18	10	1	Struttura isolante
19	10	1	Struttura isolante
20	10	1	Struttura isolante
21	10	1	Struttura isolante
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23	10	1	Struttura isolante
24	10	1	Struttura isolante
25	10	1	Struttura isolante
26	10	1	Struttura isolante
27	10	1	Struttura isolante
28	10	1	Struttura isolante
29	10	1	Struttura isolante
30	10	1	Struttura isolante
31	10	1	Struttura isolante
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48	10	1	Struttura isolante
49	10	1	Struttura isolante
50	10	1	Struttura isolante
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73	10	1	Struttura isolante
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75	10	1	Struttura isolante
76	10	1	Struttura isolante
77	10	1	Struttura isolante
78	10	1	Struttura isolante
79	10	1	Struttura isolante
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81	10	1	Struttura isolante
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90	10	1	Struttura isolante
91	10	1	Struttura isolante
92	10	1	Struttura isolante
93	10	1	Struttura isolante
94	10	1	Struttura isolante
95	10	1	Struttura isolante
96	10	1	Struttura isolante
97	10	1	Struttura isolante
98	10	1	Struttura isolante
99	10	1	Struttura isolante
100	10	1	Struttura isolante

High Resolution Mass Separator

PLOT OF (X VERSA Y) AT Z = 1.347E+01 LLL
TITLE OF GIOB INPUT: Separatore SPES caribu like 80°



High Resolution Mass Spectrometer

INFN-LNS

da Te ✨
 Rispondi Rispondi a tutti Inoltra Indesiderata
 oggetto **richiesta spettrometro magnetico di EXCIT** 15/10/2014 19:45
 a fortuna graziano ✨, giacomo cuttone ✨
 cc prete@lnl.infn.it ✨ Altre azioni

Al Presidente del MAC INFN - dott. Graziano Fortuna

cc Al direttore dei LNS -dott. Giacomo Cuttone

cc Al responsabile del progetto speciale SPES - dott. Gianfranco Prete

Oggetto: richiesta spettrometro magnetico di EXCIT per SPES

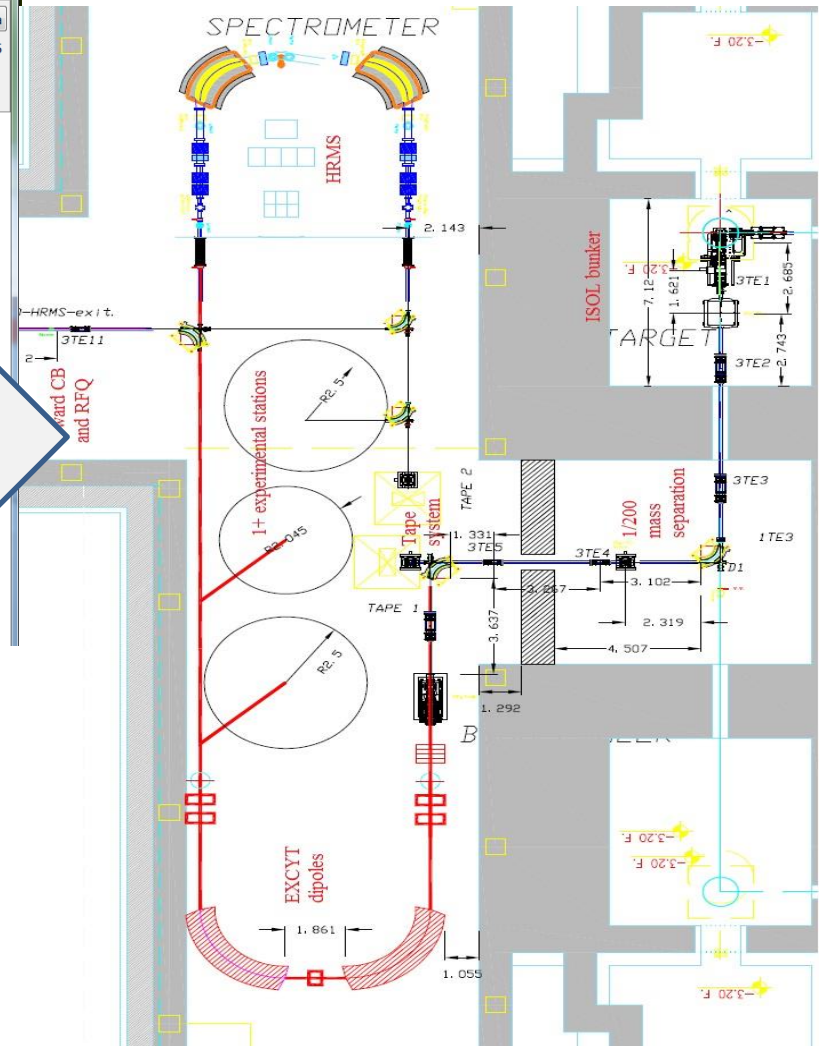
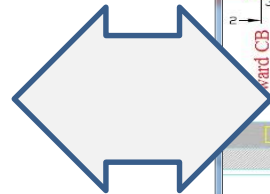
Cari colleghi,

da parte del responsabile del progetto speciale SPES trasmetto la richiesta di poter avere a disposizione presso i LNL per alcuni anni lo spettrometro magnetico di EXCIT.

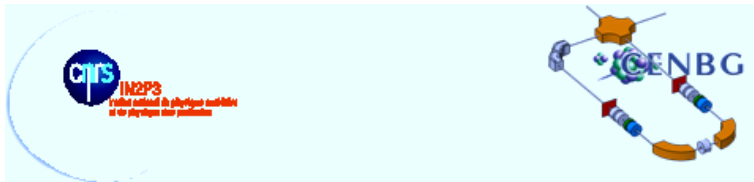
Tale strumentazione puo' essere di grande utilità come "spettrometro di primo giorno" per la selezione dei fasci del progetto SPES. Inoltre, una struttura come quella di EXCIT puo' essere vista come il primo stadio di uno spettrometro ad altissima risoluzione, da utilizzarsi nella fase di regime di SPES e da realizzarsi assieme ai LNS nell'ambito di una collaborazione internazionale che comprende SPIRAL2 e Bordeaux.

Auspico che la richiesta possa essere accolta, naturalmente nel rispetto della tempistica e della programmazione dei LNS; sarà un segno ulteriore della fraterna e proficua collaborazione già in atto fra i due laboratori.

cordiali saluti
 gianni fiorentini
 direttore LNL



CNRS-CENBG

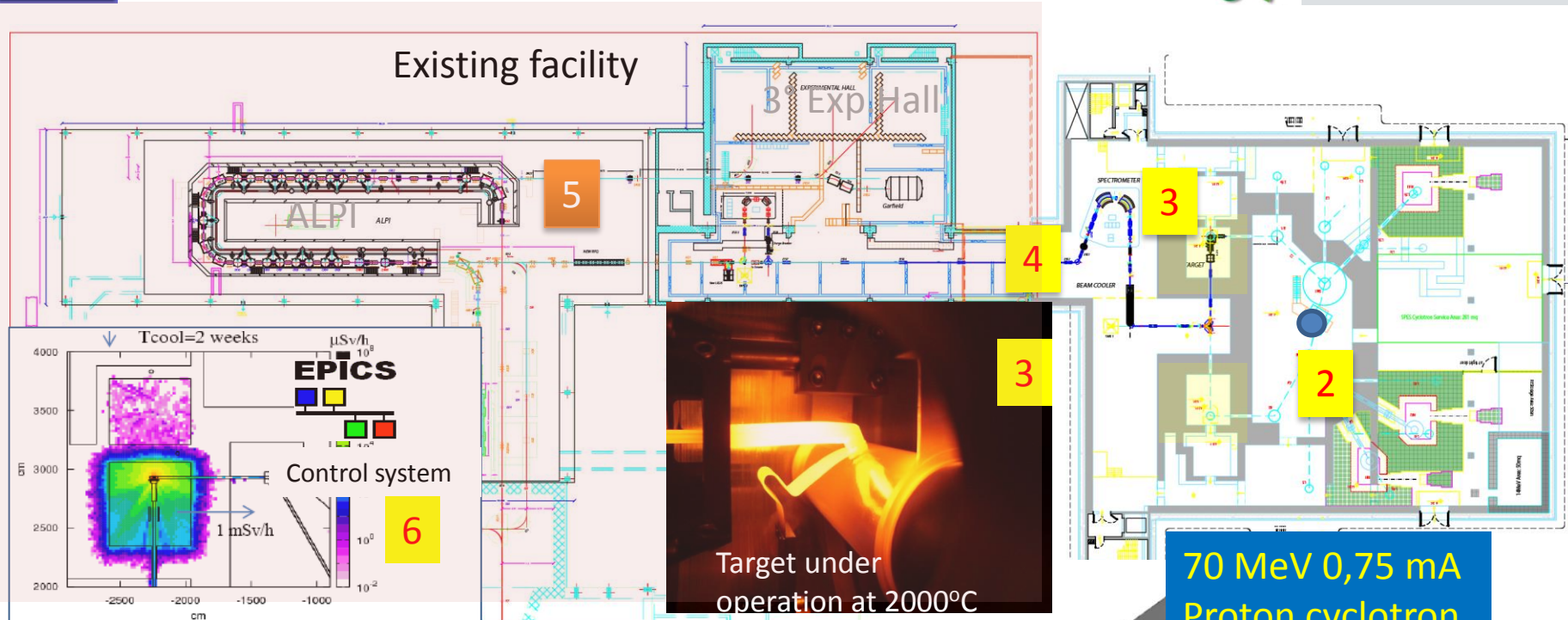


Optical design of the DESIR HRS

Teresa Kurtukian-Nieto

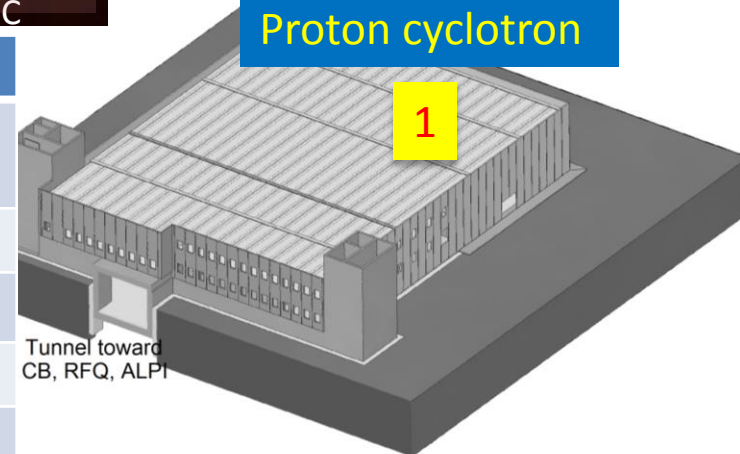
CENBG/CNRS/IN2P3-Université de Bordeaux





SPES sub-systems

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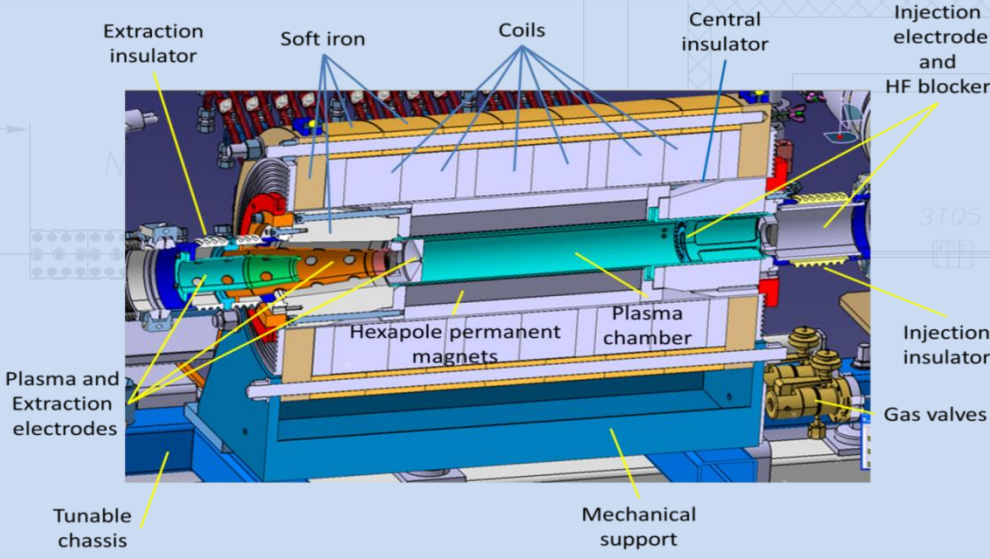
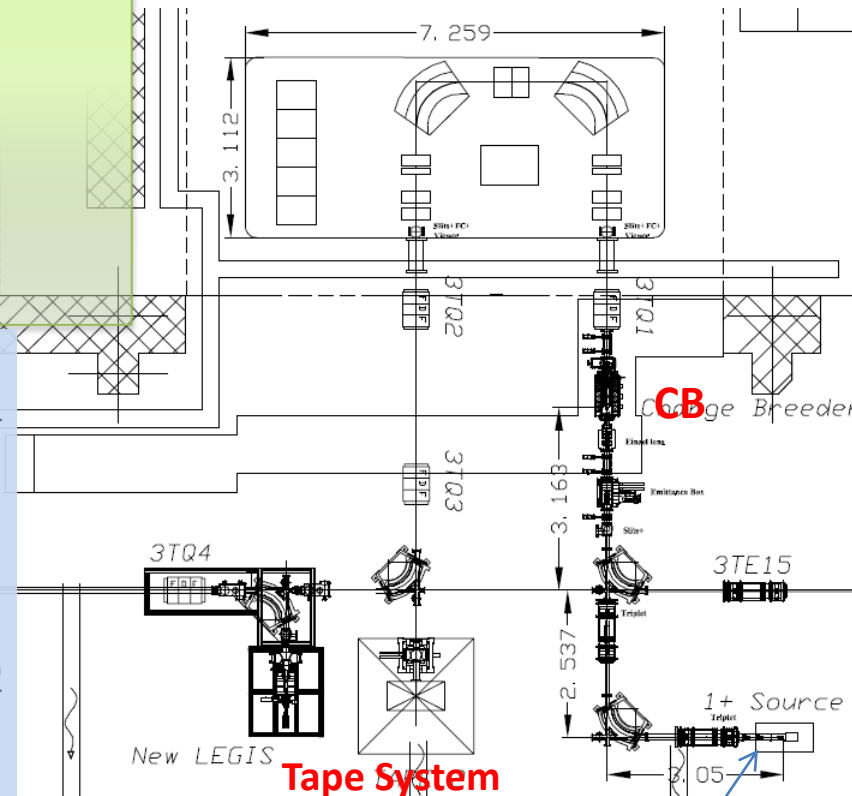
Collaboration with LPSC (Grenoble) for the SPES Charge Breeder

The development of an Upgraded PHOENIX booster is Part of a MoU in the frame of the European Associated Laboratories (LEA-Colliga) with GANIL.
(In exchange: development of SPIRAL2 n-converter by INFN)
Project and construction by LPSC_Grenoble

- 2010 Preliminary measurements
- 2011 Conceptual design and schedule definition
- 2012 Design
- 2013 Agreement definition
- 2014 Construction
- 2015 Commissioning



Mass Separator



1+ Stable Source

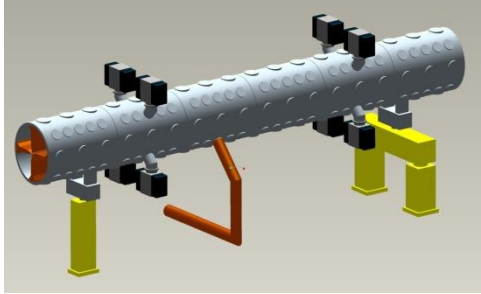
Linac Injector

Superconducting RFQs

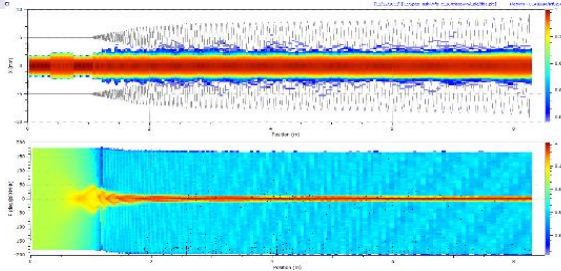


Operational since 2006

Mechanical layout of the RFQ



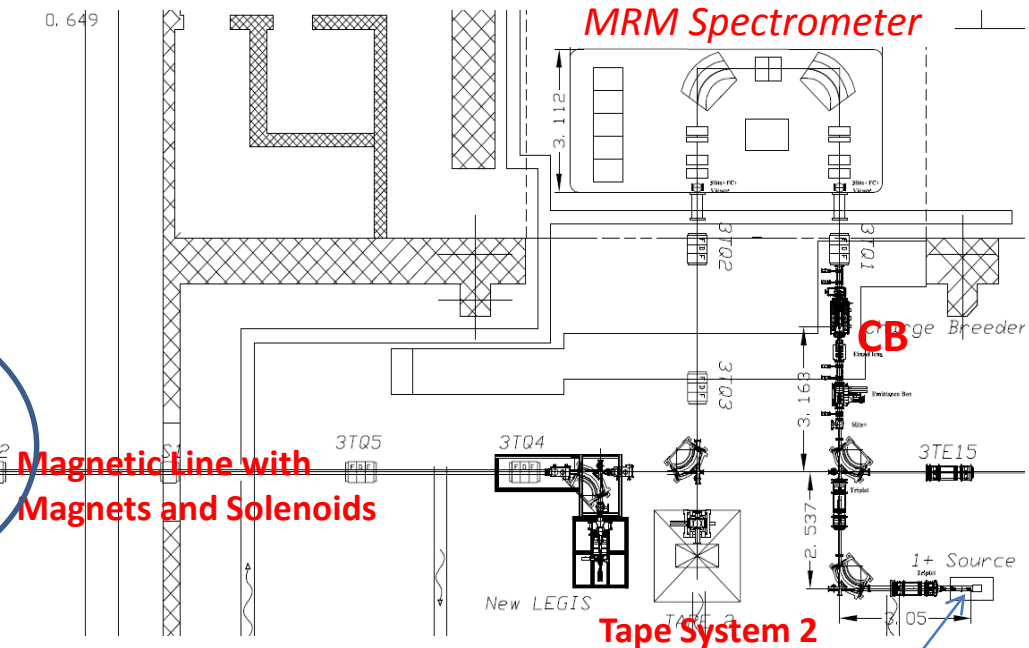
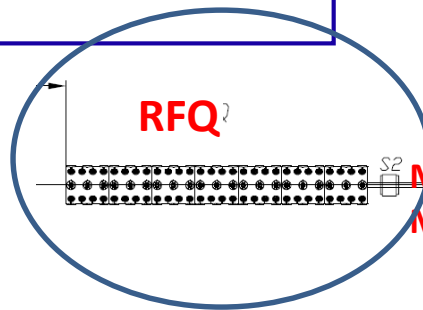
Physics design

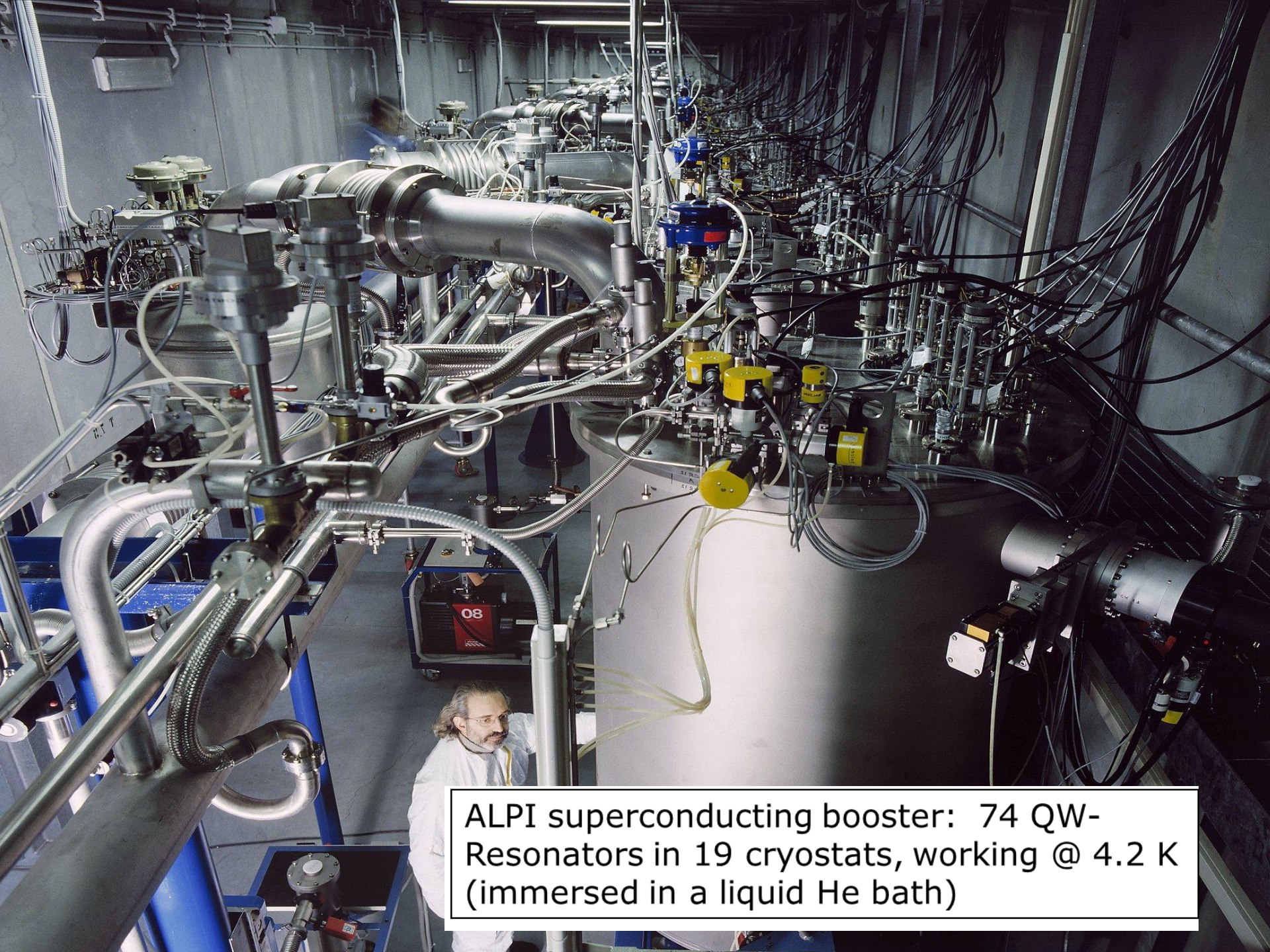


High power RF Coupler 200kW
100% duty cycle



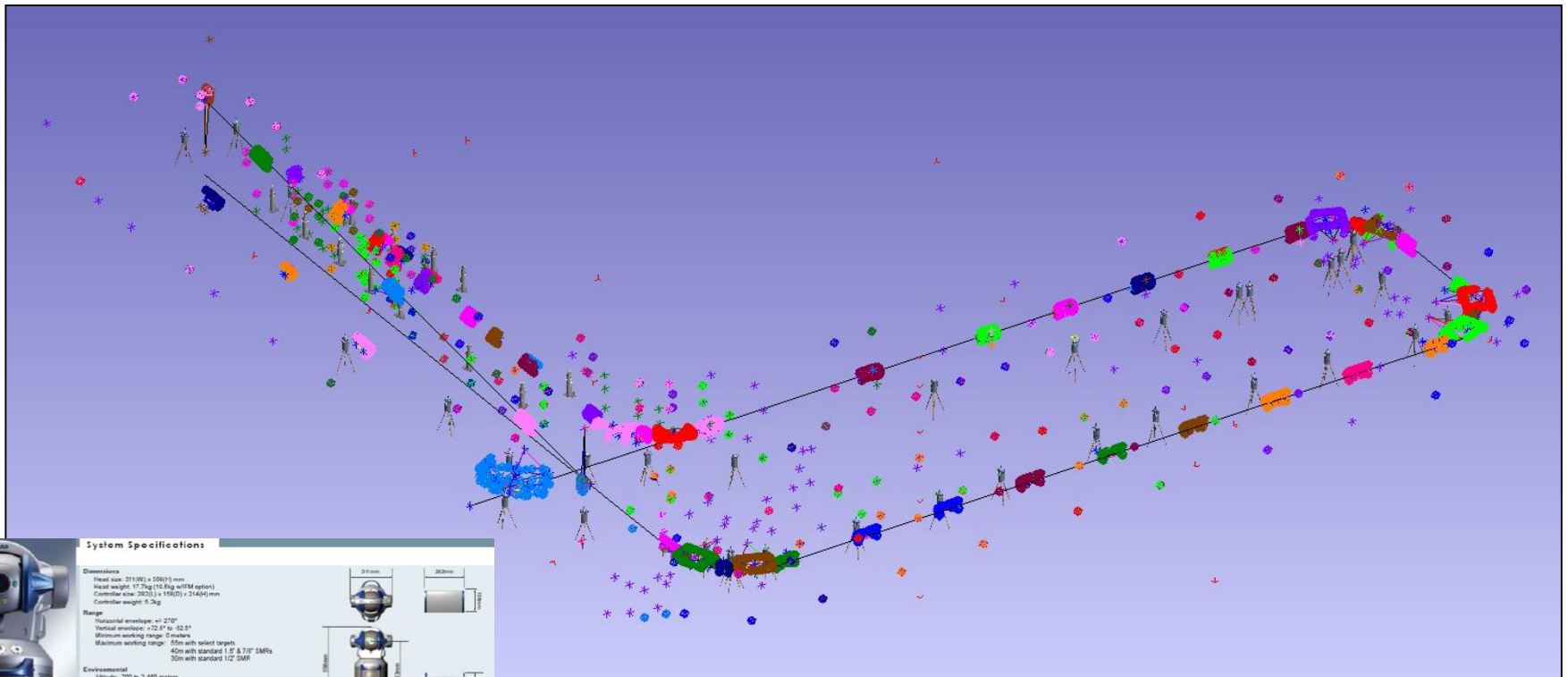
- Energy 5.7 \rightarrow 727.3 [$\beta=0.0395$] KeV/A ($A/q=7$)
- Frequency 80 MHz
- Beam transmission >95%, low RMS longitudinal emittance at output: 0.15 ns*keV/u.
- Length 695 cm (**7 modules**) intervane voltage 63.8 – 85.8 kV
- RF power (four vanes) 100 kW.
- Mechanical design and realization, taking advantage of IFMIF experience (LNL, INFN_Pd, Bo, To).





ALPI superconducting booster: 74 QW-Resonators in 19 cryostats, working @ 4.2 K (immersed in a liquid He bath)

OUTPUT OF Fiducialization OF ALL MAGNETS IN ALPI



System Specifications

Dimensions
 Head size: 211(8) x 156(1) mm
 Head weight: 17.7kg (16.6kg w/IM option)
 Controller size: 202(1) x 196(0) x 214(4) mm
 Controller weight: 6.5kg

Range
 Horizontal envelope: $\pm 270^\circ$
 Vertical envelope: $+12.8^\circ$ to -82.8°
 Minimum working range: Distance
 Maximum working range: 45m with select targets
 45m with standard 1.5" & 7/8" SMRs
 30m with standard 1/2" SMR

Environmental
 Altitude: 750 to 2,400 meters
 Humidity: 0 to 95% non-condensing
 Operating Temperature: -15°C to 50°C

Laser Emission**
 633.636 nm Laser, 1 milliwatt max/min.
 Class I Laser Product

Distance Measurement Performance

Aperture	Resolution	Accuracy
3.5mm	0.15mm	$\pm 0.15\text{mm}$
10.2mm	0.05mm	$\pm 0.05\text{mm}$
18.0mm	0.03mm	$\pm 0.03\text{mm}$
30.0mm	0.02mm	$\pm 0.02\text{mm}$

Optional Interferometric

Resolution	Accuracy
0.15mm	$\pm 0.15\text{mm}$
0.05mm	$\pm 0.05\text{mm}$
0.03mm	$\pm 0.03\text{mm}$
0.02mm	$\pm 0.02\text{mm}$

Angle Measurement Performance

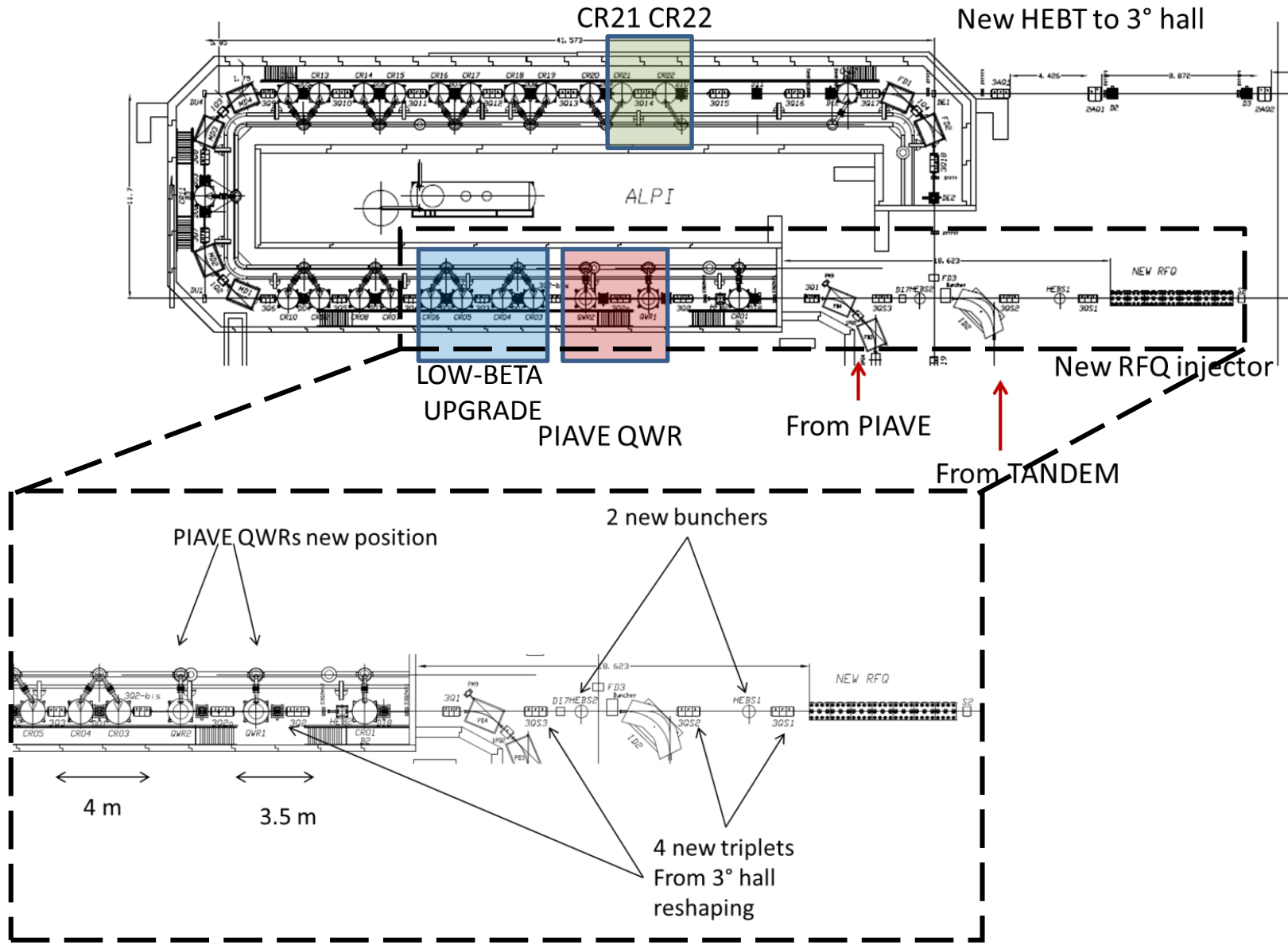
Angular accuracy (MPE)	Maximum angular velocity
$20 \mu\text{m} \pm 5 \mu\text{m/in}$	180°/sec
Precision Level Accuracy: ± 2 arcseconds	

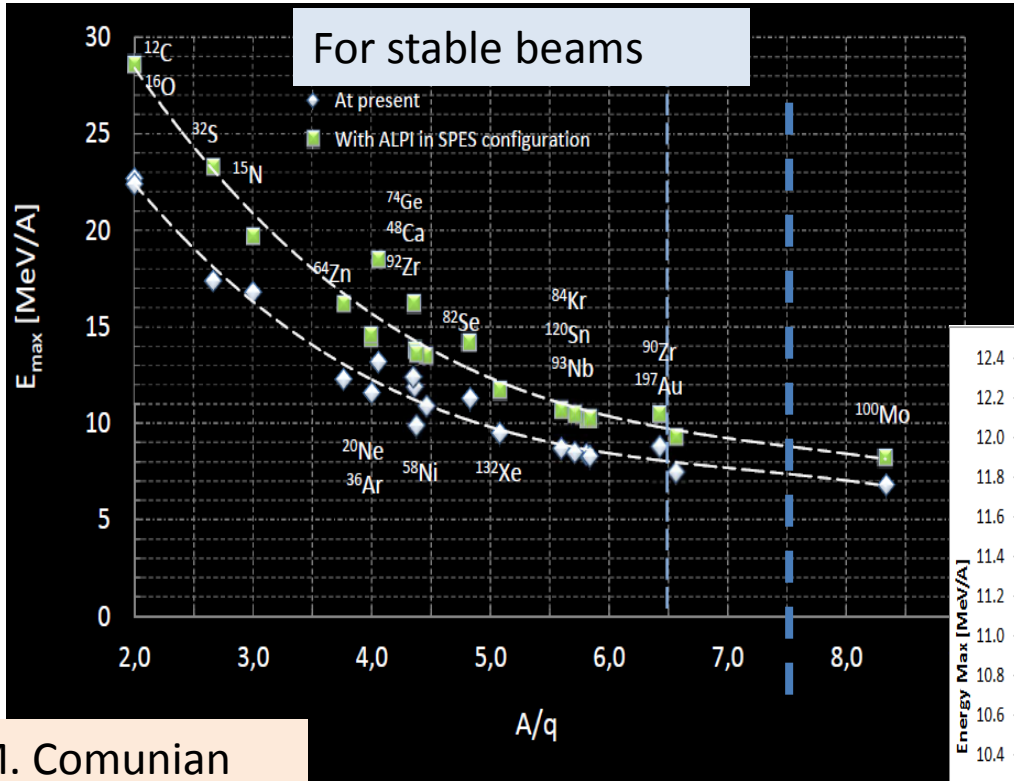
Point-to-Point Typical Accuracy***

Horizontal Scale Bar Measurement (2.3 m)	ADM (mm)	IFM (mm)
2	0.022	0.021
5	0.032	0.032
10	0.049	0.049
20	0.085	0.085
30	0.120	0.120
40	0.156	0.156
50	0.191	0.191
50°		

In-Line Distance Measurement			
Length (m)	Distance (m)	ADM (mm)	IFM (mm)
2-5	3	0.009	0.003
2-10	8	0.011	0.005
2-20	16	0.015	0.009
2-30	28	0.010	0.013
2-40	38	0.023	0.017
2-50	48	0.027	0.021

SC Resonator Improvements on ALPI

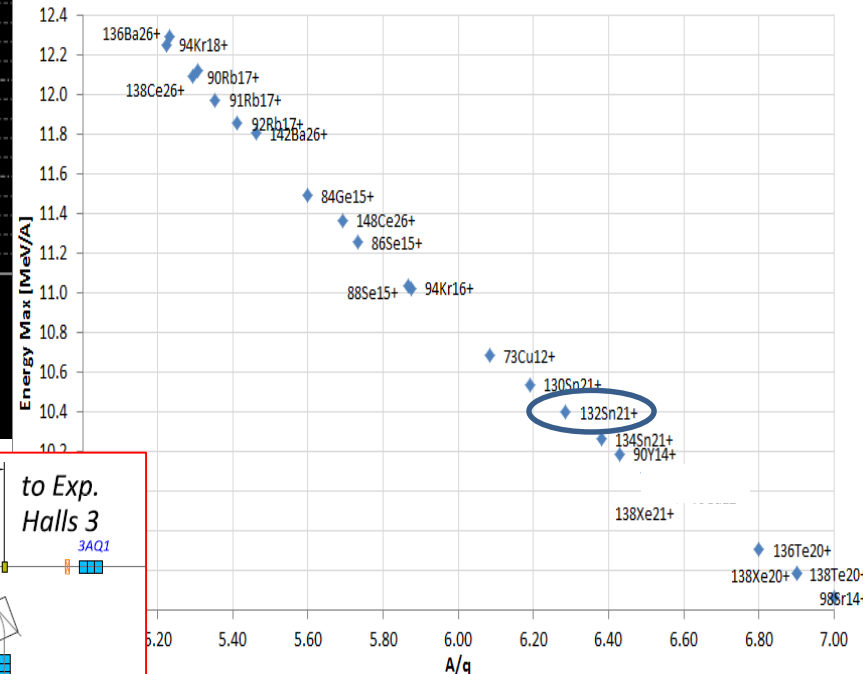




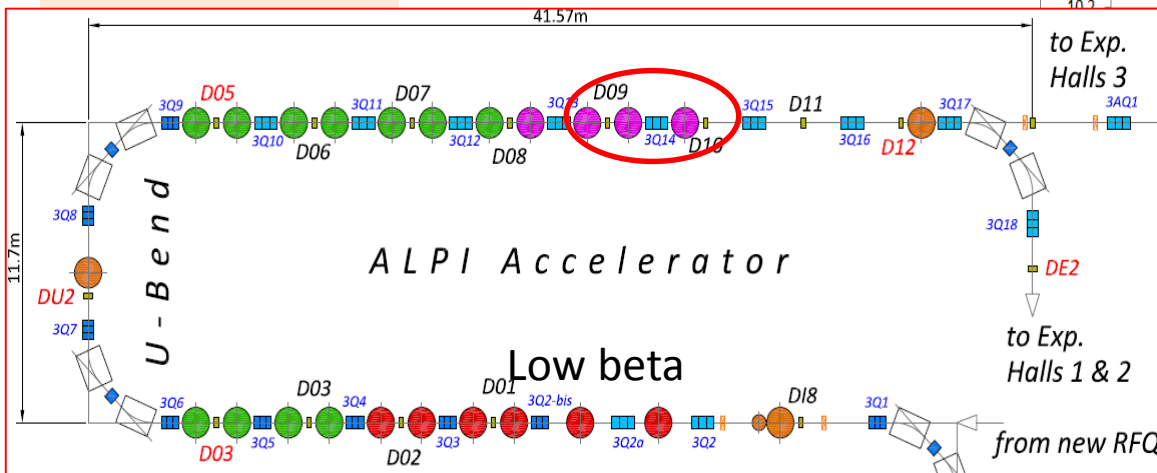
For exotic beams

Low Beta=5 MV/m, Medium Beta=4.3 MV/m, High Beta=5.5 MV/m

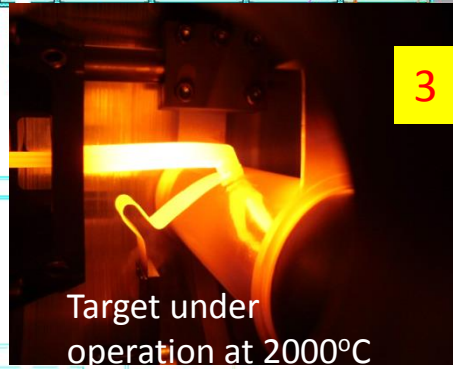
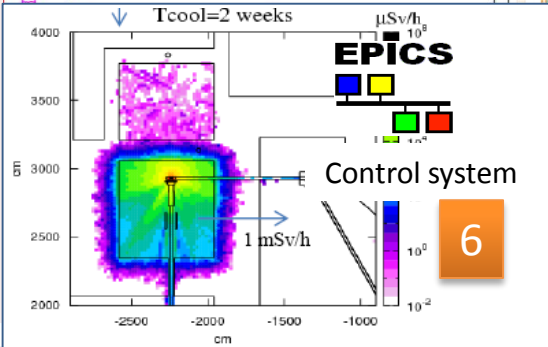
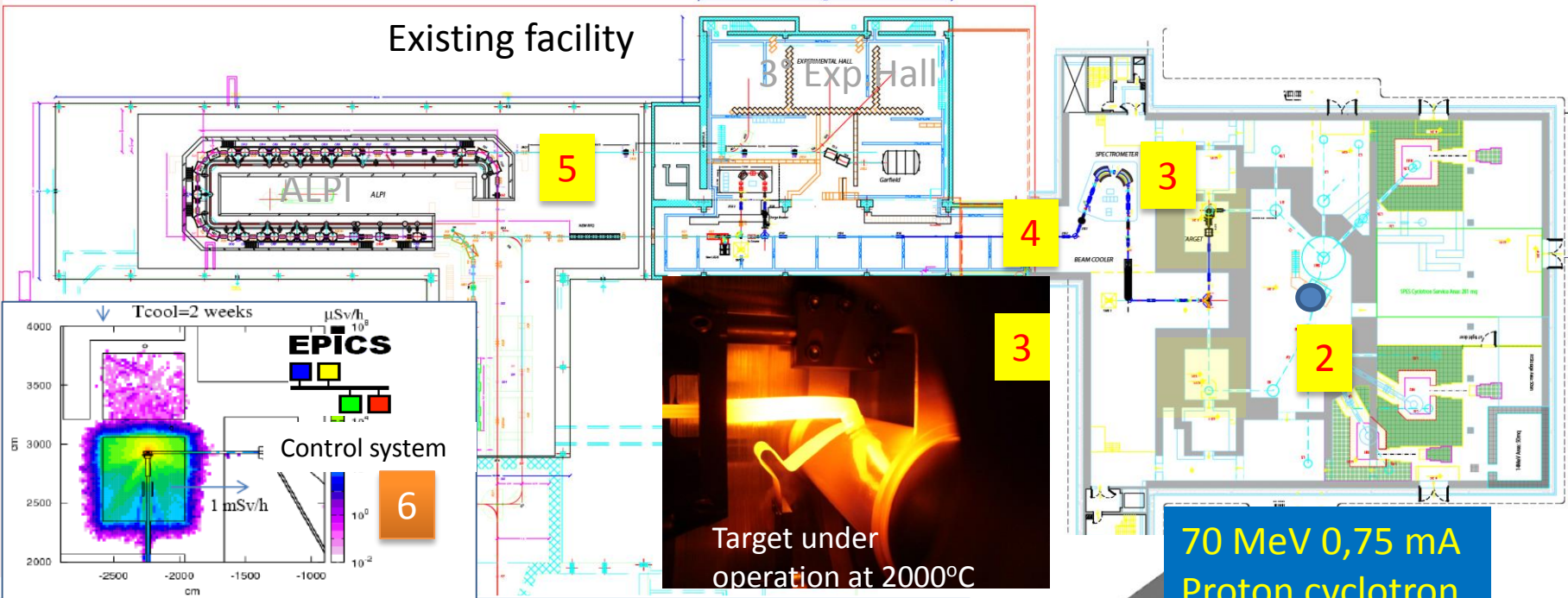
Conservative value: 2 cavities off in the calculation.



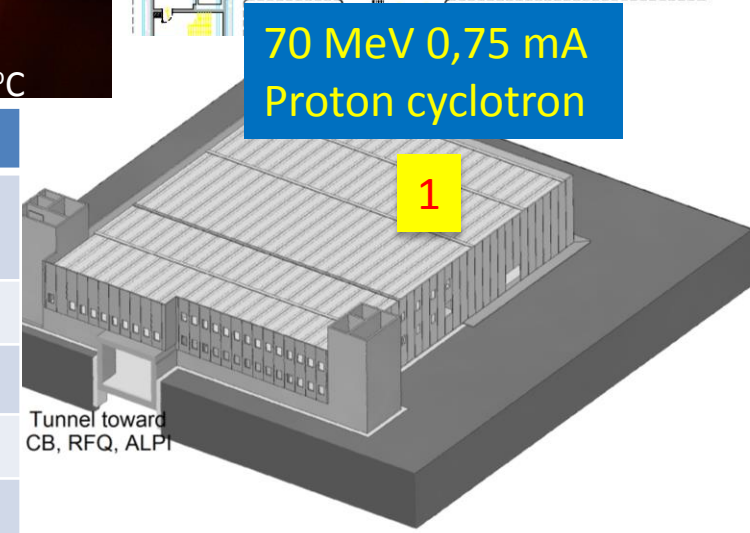
M. Comunian



Reshape and improvement of low beta cavities.
Added high beta cryostats to improve the final energy.

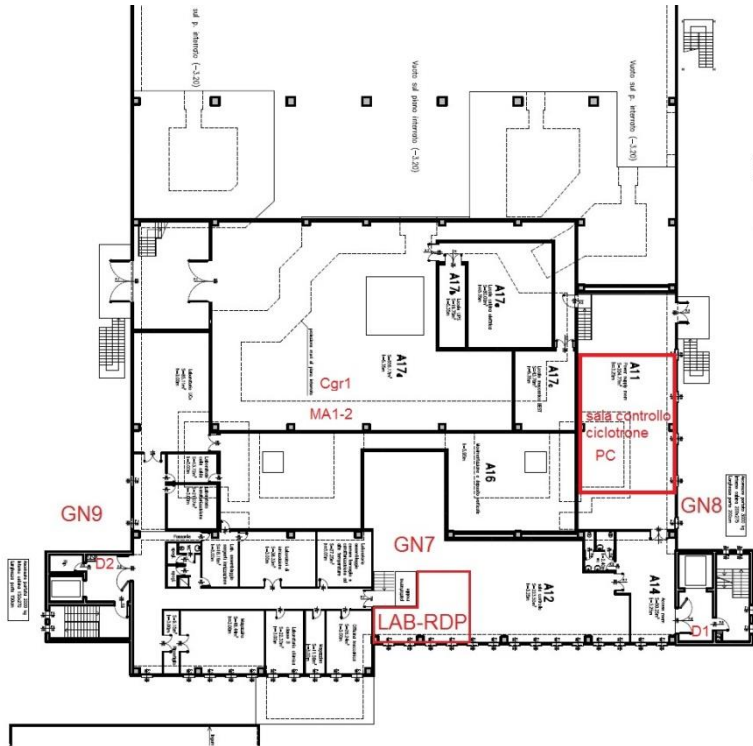
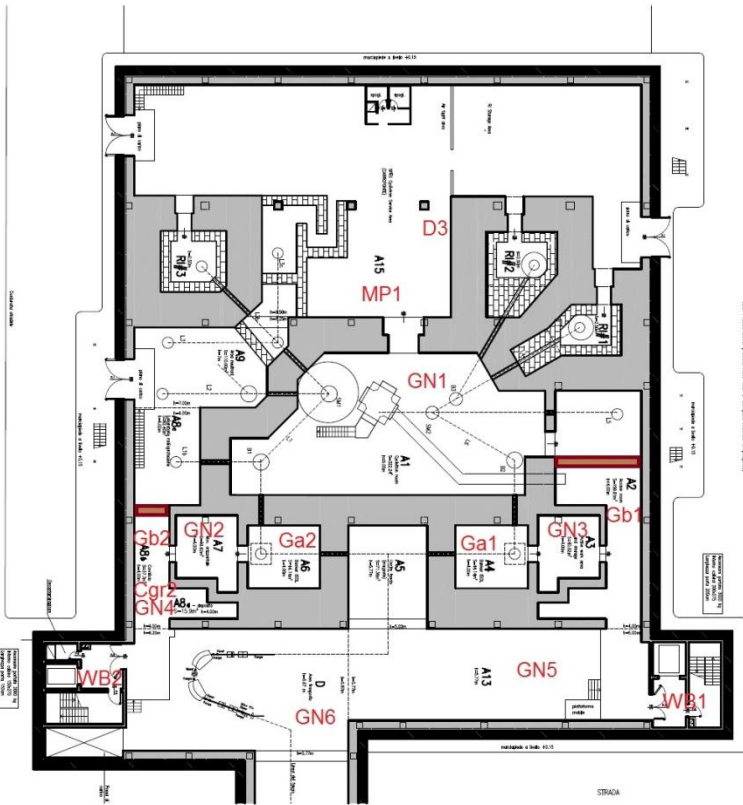
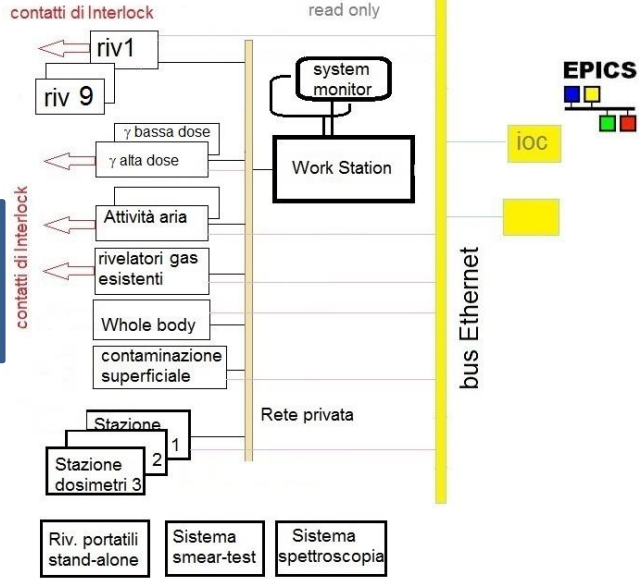


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5	Reacceleration with ALPI superconducting linac (10A MeV A=130)
6	Radioprotection, safety & controls



**70 MeV 0,75 mA
Proton cyclotron**

SPES Safety System



Actual status

Already done:

- **Building:** international bid completed, just works starting
- **Cyclotron:** on construction by BEST (Canada)
- **ISOL target:** prototype developed and under operation in lab
- **Safety & control:** authorization to the cyclotron operation just obtained, a Quality and Safety System under implementation

To be done:

- **Radioactive beam selection and transport** partially FUNDED
- **Charge breeder for increasing the charge state** FUNDED
- **RFQ for pre-acceleration** FUNDED
- **Upgrade of the ALPI superconductive Linac** partially FUNDED
- **General control system and safety** partially FUNDED

SPES 2014 may 26-28 2014 LNL

Mass Measurement via multy bounce TOF (D)

β delayed n- and γ -spectroscopy (ORNL USA)

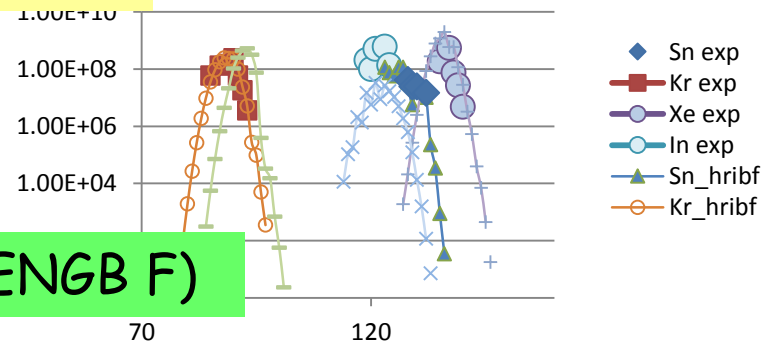
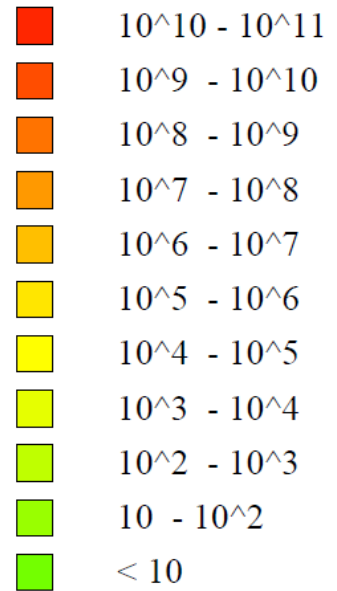
Cluster Structures via fusion evaporation with RIB (I..)

Structure studies via direct reactions (5 LOIs)

Structure studies by Safe COULEX (7 LOIs)

Collective excitations with exotic beams

β decay and r-process nucleosynthesys (CENGB F)



Decay spectroscopy techniques to study neutron-rich fission fragments at SPES

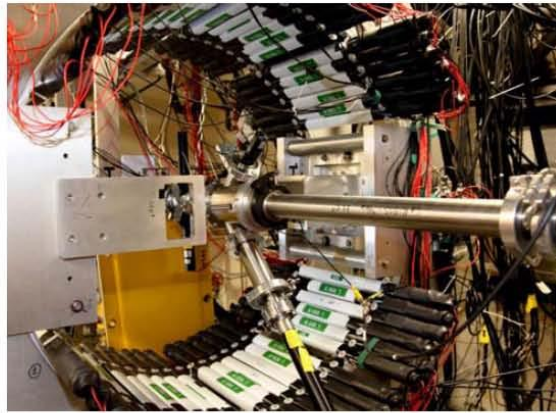
Krzysztof P. Rykaczewski, Robert Grzywacz, Carl J. Gross, Daniel W. Stracener, Yuan Liu
Oak Ridge National Laboratory, Oak Ridge, Tennessee 37831-6371, USA

in collaboration with

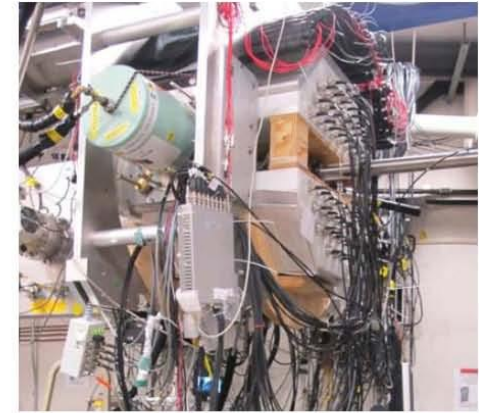
C. Mazzocchi, A. Korgul, M. Karny, K. Miernik, U. of Warsaw, Warsaw, Poland
W. Krolas, Institute of Nuclear Physics PAN, Krakow, Poland



**MTAS = Modular Total
Absorption Spectrometer**



**VANDLE = Versatile Array of
Neutron Detectors for Low Energy**



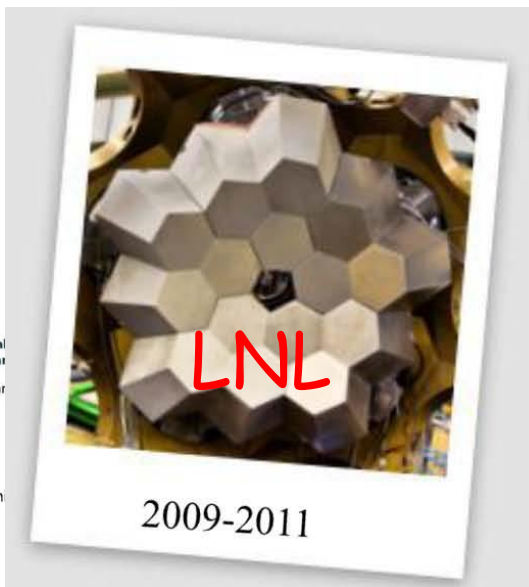
**3Hen = Helium-3 Neutron Detectors
Hybrid-3Hen = 3Hen + Clover Ge**

The physics of neutron-rich fission fragments

- nuclear structure evolution as $N \gg Z$
- spectroscopy near and above the neutron separation energy
- rapid-neutron capture half-lives and beta-delayed neutron branchings
- societal impact in better data for modeling neutron-rich environments such as nuclear reactors
- more detailed understanding of the anti-neutrino spectra from reactors

Not riaccelerated radioactive nuclear beams

To Prof. Giovanni Fiorentin
Director of LNL



Dear Gianni,

Let us first convey to you, on behalf of the AGATA Steering Committee (ASC) and AGATA Collaboration Council (ACC), the message that the full scientific community around AGATA has appreciated the interest of the LNL laboratory in the AGATA physics program and in particular in the scientific potential of the AGATA detector in combination with the exotic radioactive ion beams of the SPES facility. In view of the wide scientific program of AGATA at SPES, already envisaged by the scientific community through the presentation of 15 LOIs, the AGATA Steering Committee has agreed to install the AGATA detector at LNL-SPES for running an experimental campaign in the period 2019-2020.

Therefore the AGATA Steering Committee has decided for a commitment of the detector until 2020 (GANIL 2017-2018, LNL 2019-2020).

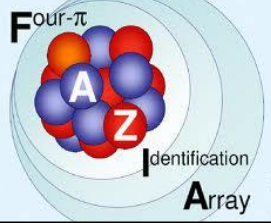
Best Regards,

G. de Angelis (ASC Chair)

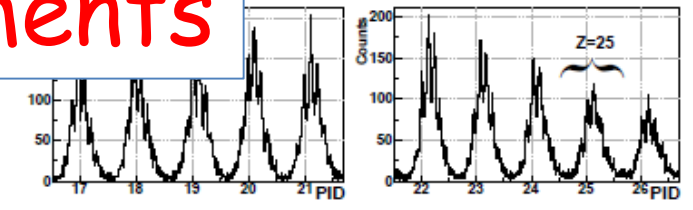
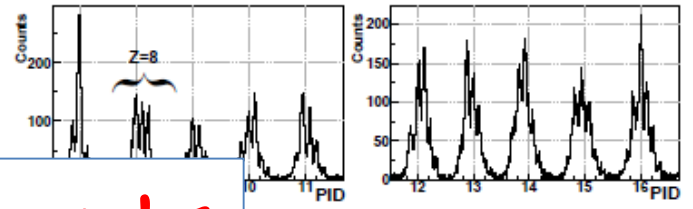
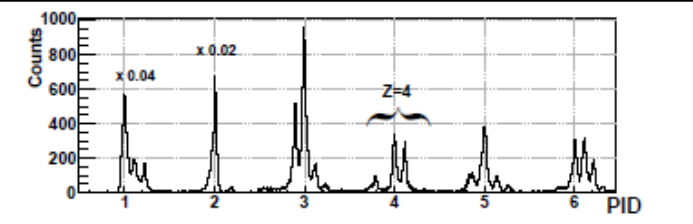
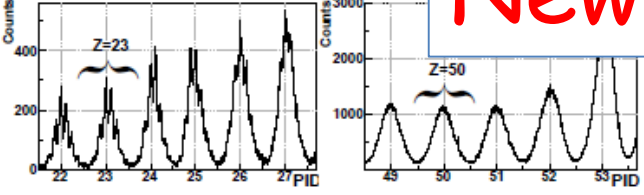
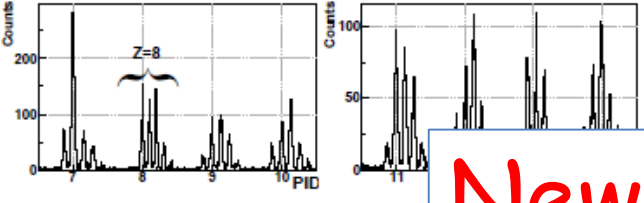
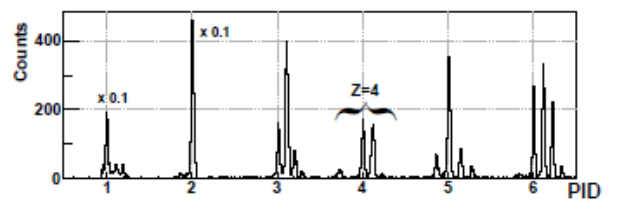
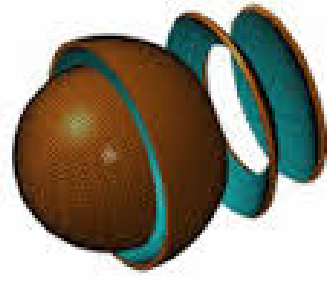
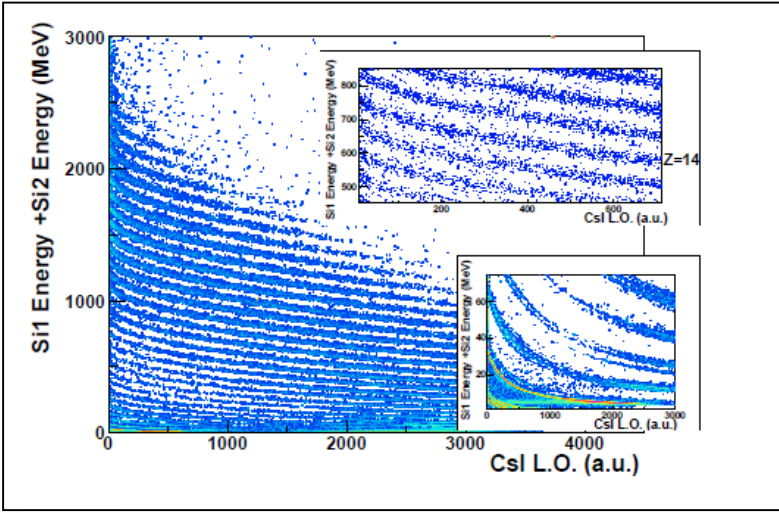
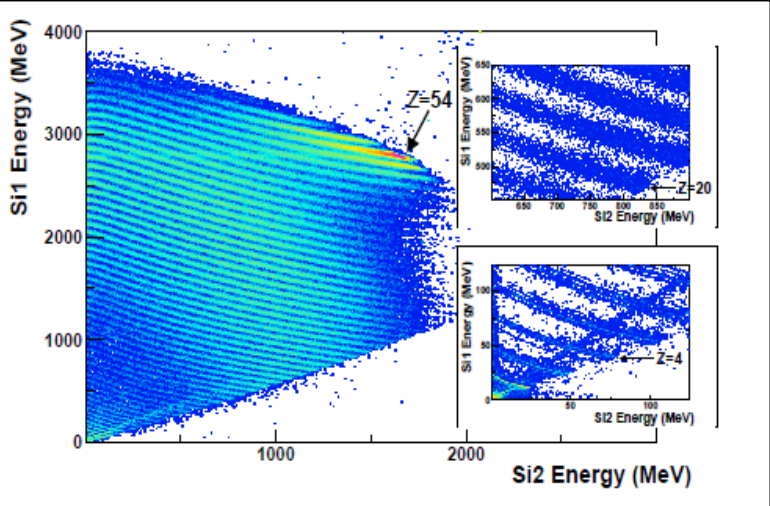
and

J. Nyberg (ACC Chair)

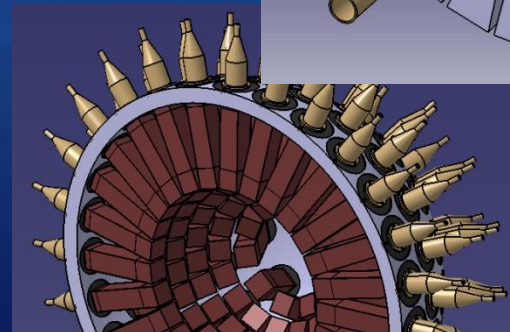
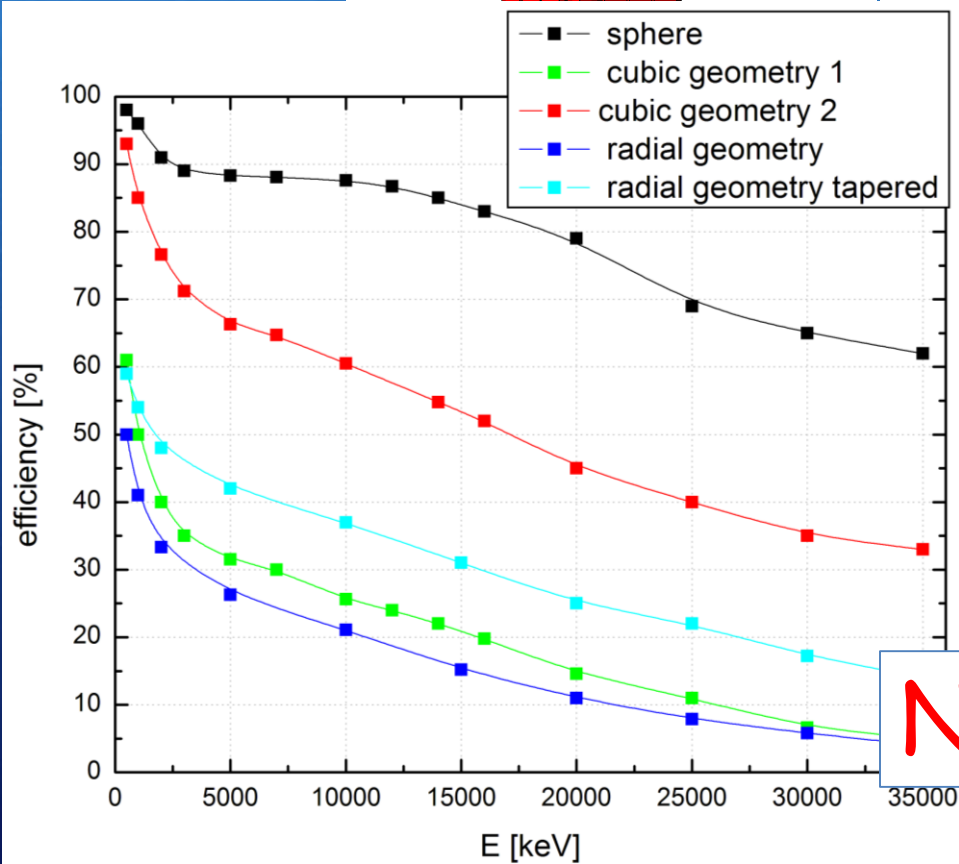
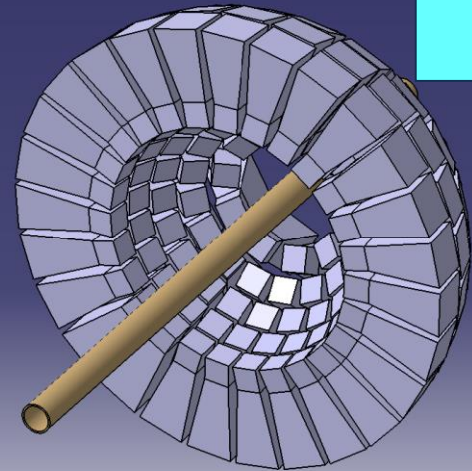
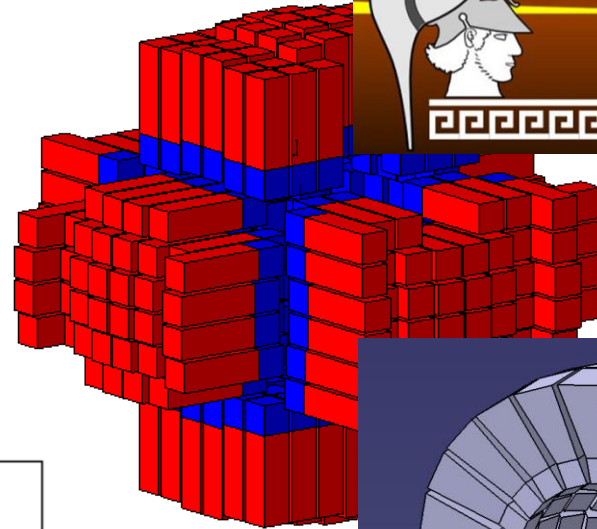
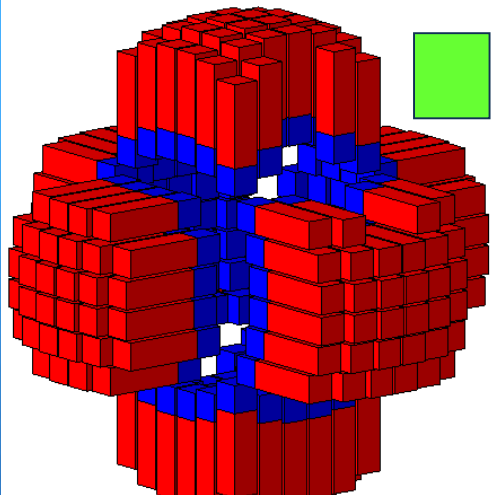
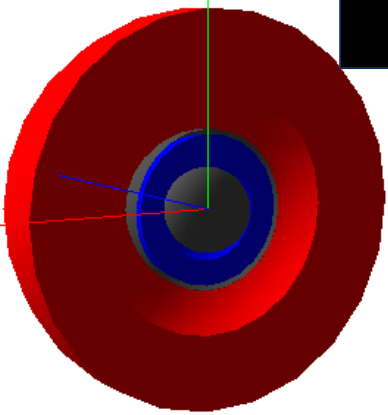
AGATA @ SPES 2019-2020



The FAZIA project



New Developments



New Developments



Thanks for attention
THANK YOU

26/11/2014
13:00