

# Experimental Research on the Reactions and Decays of Exotic Nuclei

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### **Research Activities in NRG**















Distance to center of nucleus

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- **1. Potentials of exotic nuclear systems**
- 2. Reactions with weakly-bound nuclei
- 3. 2p emissions from excited states
- 4. Decays of extremely *p*-rich nuclei



▲ Optical Model is a successful model to explain the nuclear scattering and reaction, which resembles the case of light scattered by an opaque glass sphere.

**Optical Model Potential (OMP):** 

U = V(r) + iW(r)attractive absorptive



★ phenomenological potential, independent on energy.

▲ A basic task in nuclear reaction study is to understand the nuclear interaction potential.

Cf: 1) S. Fernbach, R. Serber, and T. B. Taylor, Phys. Rev. **73**, 1352 (1949). 2) H. Feshbach, "The optical model and its justification", Ann. Rev. Nucl. Sci. **8**, 49 (1958).



# **Threshold Anomaly (TA)**





$$\Delta V(r;E) = \frac{P}{\pi} \int_0^\infty \frac{W(r;E')}{E'-E} dE'$$

Dispersion relation (results from the causality)

Cf: 1) M. A. Nagarajan, C. C. Mahaux, and G. R. Satchler, Phys. Rev. Lett. **54**, 1136 (1985). 2) C. Mahaux, H. Ngo, and G. R. Satchler, Nucl. Phys. **A449**, 354 (1986).

3) G. R. Satchler, Phys. Rep. 199, 147 (1991).

# Abnormal TA: weakly-bound nuclei

- ▲ Exotic nuclei: weakly-bound & having specified structures (cluster, halo/skin)
- ▲ Reactions: easily breakup, strongly coupling to continuum, complex mechanisms





### OMPs are usually extracted from the elastic scattering.



★ Rather difficult to extract an effective OMP at low energies.

Cf: 1) E.F. Aguilera *et al.*, PRL **84**, 5058 (2000); PRC **63**, 061603R (2001). 2) A. R. Garcia *et al.*, Phys. Rev. C **76**, 067603 (2007).



### **Transfer Method**



<sup>63</sup>Cu(<sup>7</sup>Li, <u><sup>6</sup>He</u>)<sup>64</sup>Zn: Phys. Rev. C **95**, 034616 (2017).

### Experiments: <sup>208</sup>Pb(<sup>7</sup>Li,<sup>6</sup>He)<sup>209</sup>Bi

Two experiments have been done at HI-13 tandem accelerator @ CIAE Exp1:  $E_{\text{beam}} = 42.55$ , 37.55, 32.55, 28.55, 25.67 MeV – high energies [2004.8] Exp2:  $E_{\text{beam}} = 28.55$ , 25.67, 24.3, 21.2 MeV – low energies [2016.4] \* Angular distributions of both elastic scattering and transfer were measured.





### **Results:** OMPs of <sup>6</sup>He+<sup>209</sup>Bi



- ★ OMPs of the <sup>6</sup>He+<sup>209</sup>Bi system are determined precisely for the first time;
- ★ The decreasing trend in the imaginary part is observed, and the threshold energy is about 13.73 MeV (~0.68V<sub>B</sub>);
- ★ The behavior of real part looks normal, i.e. like a bell shape around the barrier;
- ★ The dispersion relation does
   NOT hold in this system.

L. Yang, C.J. Lin\*, H.M. Jia et al., Phys. Rev. Lett. **119**, 042503 (2017); Phys. Rev. C **96**, 044615 (2017).



### Discussions

- **★** Dispersion relation results from causality, connecting real and imaginary part;
- \* Any wave/particle should follow this relation when it passes through a media;
- **★** The dispersion relation is **not** applicable for exotic nuclear systems.

**Possible reasons:** 

- Causality → dispersion relation stable systems: causality ↔ analyticity
- Cauchy integration infinity poles (breakup) & off-axis (multi-process)
- Negative Index of Refraction causality based criteria must be used with care [Phys. Rev. Lett. 101, 167401 (2008).]
- Locality vs. non-locality equivalent local potential in Schrödinger equation





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### **1. Potentials of exotic nuclear systems**

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# **Reactions with Exotic Nuclei**

**RIBs experiments** 

Elastic scattering
 3-body, 4-body
 CDCC ...

Fusion/Reaction
TF = ICF + CF ...

Breakup/transfer
Effects & mechanisms



[1] L. F. Canto, P. R. S. Gomes, R. Donangelo et al., Phys. Rep. 424, 1 (2006).
[2] L. F. Canto, P. R. S. Gomes, R. Donangelo et al., Phys. Rep. 596, 1 (2015).
[3] B. B. Back, H. Esbensen, C. L. Jiang and K. E. Rehm, Rev. Mod. Phys. 86, 317 (2014).



### **Reaction mechanism**

★ How to identify the different reaction process?





### **Experiments**

### **★** Complete-kinematics measurement ; **★** Reactions induced by <sup>7</sup>Be, <sup>8</sup>B, <sup>17</sup>F ...



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### Preliminary Results: 17F+58Ni



# Preliminary Results: <sup>17</sup>F+<sup>58</sup>Ni



[Cf. Lei Jin & A. M. Moro, PRL 122, 042503 (2019)]

### **Preliminary conclusions:**

- The non-elastic breakups are dominant;
- Fusions are suppressed at energies above the barrier but enhanced below the barrier.



### Discussions

★ Exclusive breakup (<sup>16</sup>O-*p*)

Our result: **σ ~ 1.2 mb** @ **63 MeV**;

Liang's result:  $\sigma \sim 15.6 \text{ mb} @ 170 \text{ MeV}$ .

[J.F. Liang et al., PLB 681, 22 (2009).]



Why are the breakup cross sections so low?

- Screen effects due to the dynamic polarization?
- Transfers are dominant?
- ...

<sup>8</sup>B+<sup>120</sup>Sn experiment will be performed at CNS/RIKEN (2 -16 Apr., 2019)

<sup>17</sup>F ( $S_p = 0.601$  MeV), <sup>8</sup>B ( $S_p = 0.136$  MeV)



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# Exotic decays of *p*-rich nuclei



 $\beta p$ ,  $\beta 2 p$ ,  $\beta 3 p$ ,  $\beta p \gamma$ ,  $\beta \gamma p$ ,  $\beta \alpha$ ,  $\beta 2 \alpha$ ,  $\beta \alpha p$ ,  $\beta p \alpha$ ,  $\beta p 2 \alpha$ ,  $\beta n$ ,  $\beta 2 n$ ,  $\beta 3 n$ ,  $\beta d$ ,  $\beta t$ ,  $\beta F$ ...

- Structures of *p*-rich nuclei close to/beyond the drip-line
- Effective interaction pairing, isospin non-conserving (INC), three-body force
- Initial state interaction (ISI), final state interaction (FSI), quantum entanglement
- Nuclear astrophysics  $(p, \gamma)$ ,  $(2p, \gamma)$ ,  $(\alpha, \gamma)$  ... processes



### **Overview of our research**

													<sup>40</sup> Sc	<sup>41</sup> Sc
		Sta	ble	Nuc	lide			20	<sup>35</sup> Ca	<sup>36</sup> Ca	<sup>37</sup> Ca	<sup>38</sup> Ca	<sup>39</sup> Ca	<sup>40</sup> Ca
		Implantation								<sup>35</sup> K	<sup>36</sup> K	<sup>37</sup> K	<sup>38</sup> K	<sup>39</sup> K
		In-flight				18	<sup>31</sup> Ar	<sup>32</sup> Ar	<sup>33</sup> Ar	<sup>34</sup> Ar	<sup>35</sup> Ar	<sup>36</sup> Ar	<sup>37</sup> Ar	<sup>38</sup> Ar
								<sup>31</sup> Cl	<sup>32</sup> CI	<sup>33</sup> CI	<sup>34</sup> CI	<sup>35</sup> CI	<sup>36</sup> CI	<sup>37</sup> Cl
				16	<sup>27</sup> S	<sup>28</sup> S	<sup>29</sup> S	<sup>30</sup> S	<sup>31</sup> S	<sup>32</sup> S	<sup>33</sup> S	<sup>34</sup> S	<sup>35</sup> S	<sup>36</sup> S
	<sup>26</sup> P <sup>27</sup> P <sup>28</sup> P						<sup>28</sup> P	<sup>29</sup> P	<sup>30</sup> P	<sup>31</sup> P	<sup>32</sup> P	<sup>33</sup> P	<sup>34</sup> P	<sup>35</sup> P
	14	<sup>22</sup> Si	<sup>23</sup> Si	<sup>24</sup> Si	<sup>25</sup> Si	<sup>26</sup> Si	<sup>27</sup> Si	<sup>28</sup> Si	<sup>29</sup> Si	<sup>30</sup> Si	<sup>31</sup> Si	<sup>32</sup> Si	<sup>33</sup> Si	<sup>34</sup> Si
	•		<sup>22</sup> AI	<sup>23</sup> AI	<sup>24</sup> AI	<sup>25</sup> AI	<sup>26</sup> AI	<sup>27</sup> AI	<sup>28</sup> AI	<sup>29</sup> AI	<sup>30</sup> AI	<sup>31</sup> AI	<sup>32</sup> AI	<sup>33</sup> AI
	12	<sup>20</sup> Mg	<sup>21</sup> Mg	<sup>22</sup> Mg	<sup>23</sup> Mg	<sup>24</sup> Mg	<sup>25</sup> Mg	<sup>26</sup> Mg	<sup>27</sup> Mg	<sup>28</sup> Mg	<sup>29</sup> Mg	<sup>30</sup> Mg	<sup>31</sup> Mg	<sup>32</sup> Mg
	•		<sup>20</sup> Na	<sup>21</sup> Na	<sup>22</sup> Na	<sup>23</sup> Na	<sup>24</sup> Na	<sup>25</sup> Na	<sup>26</sup> Na	<sup>27</sup> Na	<sup>28</sup> Na	<sup>29</sup> Na	<sup>30</sup> Na	<sup>31</sup> Na
10	<sup>17</sup> Ne	<sup>18</sup> Ne	<sup>19</sup> Ne	<sup>20</sup> Ne	<sup>21</sup> Ne	<sup>22</sup> Ne	<sup>23</sup> Ne	<sup>24</sup> Ne	<sup>25</sup> Ne	<sup>26</sup> Ne	<sup>27</sup> Ne	<sup>28</sup> Ne	<sup>29</sup> Ne	<sup>30</sup> Ne
		<sup>17</sup> F	<sup>18</sup> F	<sup>19</sup> F	<sup>20</sup> F	<sup>21</sup> F	<sup>22</sup> F	<sup>23</sup> F	<sup>24</sup> F	<sup>25</sup> F	<sup>26</sup> F	<sup>27</sup> F		<sup>29</sup> F

- Started from 2004
- ♣ RIBLL@HIRFL (Lanzhou)

### ♠ In-flight decay

(Ex. states 2p emissions) <sup>28,29</sup>S/<sup>27,28</sup>P;

<sup>17,18</sup>Ne.

- ♠ Implantation decay
  - (G.S. βp, β2p...) <sup>36,37</sup>Ca; <sup>27</sup>S/<sup>26</sup>P/<sup>25</sup>Si; <sup>22</sup>Si/<sup>20</sup>Mg; <sup>23</sup>Si/<sup>22</sup>Al/<sup>21</sup>Mg; <sup>24</sup>Si/<sup>23</sup>Al.



# In-flight decays

### **2***p* emissions from high-lying excited states and related topics



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### Exp. setup 1

**RIBLL** Experimental Setup 2007



# Covered ±10° 255mm

#### Note:

**Collimators:** PPAC1: φ30 mm; PPAC2: φ20 mm. ΔE: 300μm Si ΔE detector, combined with TOF (I <sup>197</sup>Au: target, 200-250 μm, φ28 mm. D1,D2,D3: 300μm Si, 48mm×48mm. D4: 1000μm Si detector with 4 segments, 50mm× X1,Y1,X2,Y2: 300μm Si strip detectors, each of 2 CSI: CSI(TI)+PIN detectors, 20 mm length, total C

Complete-kinematics measurements



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### Exp. setup 2

SSSD3,4

р



Secondary target: <sup>197</sup>Au, 100 µm SD: Silicon detectors, 325, 1000 µm

SSSD: Single sided Silicon Strip Detec with 2 mm in the width and 0.1 the construction of the particle tr CsI(Tl) array:  $6 \times 6$  lattices, each 15 through **PIN** photodiode. *Complete-kinematics* measurements







### 2p-decay modes



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### 2*p* emission: <sup>29</sup>S/ <sup>28</sup>P



#### Diproton emissions were observed in <sup>29</sup>S but not in <sup>28</sup>P.

<sup>29</sup>S: C.J. Lin, X. X. Xu, H. M. Jia *et al.*, PRC **80**, 014310 (2009);
<sup>28</sup>P: X. X. Xu, C.J. Lin, H.M. Jia *et al.*, PRC **81**, 054317 (2010).

### 2*p* emission: <sup>28</sup>S/<sup>27</sup>P

**\star** Diproton emission is enhanced by 2*p* halo-like states.



X.X. Xu, C.J. Lin\* et al., Phys. Lett. B 727, 126 (2013).

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### 2p emission & 2p halo

#### 2p halo/skin in proton-rich S isotopes





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# BCS/BEC crossover: 17,18 Ne



C.J. Lin\*, X.X. Xu, H.M. Jia et al., JPS Conf. Proc. 1, 013026 (2014).



### Discussions

### **Question: How to describe** *2p* **emission in precise?**





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### **Implantation decays**

 $\beta$ -decay spectroscopy of nuclei close to the proton drip line



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### Exp. setup1



A detector array for 2*p*-decay study by **implantation** method for lifetime > 10 μs

1p efficiency: 66%; 2p efficiency: 20%





### Exp. setup2



2p efficiency: 20%

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Results 1: <sup>22</sup>Si/<sup>20</sup>Mg

Primary beam: <sup>28</sup>Si, 75.3 MeV/u @ 40 enA. <sup>35</sup>Ca <sup>36</sup>Ca  $\beta p$  precursors 10<sup>4</sup> 35K <sup>22</sup>Si: 4×10<sup>-3</sup> pps @ 8×10<sup>-4</sup>% 3000 33 Ar <sup>34</sup>Ar 32 A/ <sup>20</sup>Mg: 0.72 pps @ 0.15% 33CI 31 CI 32CI 2500 10<sup>3</sup> N=827S <sup>31</sup>S <sup>32</sup>S 28S 29S 30S <sup>18</sup>Ne 26p 28p 29p <sup>30</sup>P <sup>31</sup>P 27P 2000 17 F DE1 23Si 27Si  $T_{z} = -3$ 24Si 25Si <sup>26</sup>Si <sup>28</sup>Si 29Si <sup>30</sup>Si 22Si 10<sup>2</sup> 16**0** 1500 23 AI 22AI 24 AI 25AI <sup>26</sup>AI 27AI 14**6**  $T_z = -2$ <sup>24</sup>Mg <sup>21</sup>Mg <sup>22</sup>Mg <sup>23</sup>Mg <sup>24</sup>Mg <sup>25</sup>Mg <sup>26</sup>Mg 1000 10 <sup>20</sup>Na <sup>21</sup>Na <sup>22</sup>Na <sup>23</sup>Na <sup>18</sup>Ne <sup>19</sup>Ne <sup>20</sup>Ne <sup>21</sup>Ne <sup>22</sup>Ne Ne 500 Atta Darieta de Constante -300 -250 -200 -700 -650 -600-550 -500-450 -400 -350 T1-T2

PID of the secondary beam

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Doolz	Energy	BR	Decay		
ГСак	(keV)	(%)	Mode		
1	230(50)	<b>2.9(10)</b>	2p ?		
2	680(50)	6.8(14)	βp		
3	1710(50)	1.9(7)	βp		
4	1950(50)	52.0(74)	βp		
5	2110(50)	10.9(21)	βp		
6	2180(50)	6.5(15)	βp		
7	2330(50)	5.1(13)	βp		
8	3550(50)	2.5(9)	βp		
9	5600(70)	0.7(3)	<i>β</i> 2 <i>p</i>		

#### ★ Mass of <sup>22</sup>Si

•  $\Delta(^{22}\text{Si}) = \Delta(^{22}\text{Al IAS}) + \Delta E_{\text{C}} - \Delta_{n\text{H}}$   $\rightarrow S_{2p} = -108 \pm 125 \text{ keV};$ •  $\Delta(^{22}\text{Si}) = \Delta(^{22}\text{O}) - 2b(A,T)T_{\text{Z}}$  $\rightarrow S_{2p} = -15 \text{ keV}$ 

The first experimental mass data. The first  $\beta 2p$  precursor found in Asian Lab.

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# Discussions on <sup>22</sup>Si/<sup>20</sup>Mg

### Solution Mirror asymmetry $\rightarrow$ INC interaction asymmetry parameter: $\delta = \frac{ft^+}{ft^-} - 1$

20	<sup>0</sup> Mg→ <sup>20</sup> Na				$^{20}O \rightarrow ^{20}F$		
<sup>20</sup> Na $E^*$ (keV)	br (%)	log ft	$J^{\pi}$	$^{20}$ F $E^*$ (keV)	br (%)	log ft	δ
983.9(22)	66.9(46)	3.80(4)	1+	1056.848(4)	99.973(3)	3.740(6)	0.148(107)
2998(13)	8.6(7)	4.15(4)	1+	3488.54(10)	0.027(3)	3.65(6)	2.16(53)
	<sup>22</sup> Si→ <sup>22</sup> Al				$^{22}O \rightarrow ^{22}F$		
$^{22}$ Al $E^*$ (keV)	br (%)	log ft	$J^{\pi}$	$^{22}$ F $E^*$ (keV)	br (%)	log ft	$\delta$
1170(50)	5.1(3)	5.10(5)	1+	1625	29(4)	4.6(1)	2.16(82)
2400(50)	60.6(65)	3.79(7)	1+	2572	68(6)	3.8(1)	-0.02(28)

### Three-Body Force PRL110,022502(2013).

		$S_p$		$S_{2p}$			
Nucleus	Expt.	NI	V + 3N	Expt. N		N + 3N	
N = 8	[IMME]	sd	$sdf_{7/2}p_{3/2}$	[IMME]	sd	$sdf_{7/2}p_{3/2}$	
<sup>18</sup> Ne	3.92	4.05	3.76	4.52	4.67	4.17	
<sup>19</sup> Na	-0.32	-0.32	-0.26	3.60	3.73	3.50	
$^{20}$ Mg	2.66	2.83	2.98	2.34	2.51	2.72	
$^{21}Al$	[-1.34]	-2.52	-1.83	[1.45]	0.30	1.15	
<sup>22</sup> Si	[1.35]	0.90	1.71	[0.01]	-1.63	-0.12	



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### Results 2: <sup>27</sup>S



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### Daughter: <sup>27</sup>P

 $\beta p \& \beta \gamma$  were measured simultaneously for the first time.



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### The Galactic <sup>26</sup>Al puzzle



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# Thermonuclear <sup>26</sup>Si(*p*,γ)<sup>27</sup>P Rate



Comparison of the calculated thermonuclear reaction rates from the  $3/2^+$  resonance contribution.

L.J. Sun *et al.*, arXiv:1809.02980v1; arXiv:1809.02987v1.

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### Collaborators

### Thanks to all the collaborators

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# Thank you for your attention.

