Shell evolution along the Sn isotopes

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Introduction

- What is interesting in the ¹⁰⁰Sn region?
 - 1. Aspect of the heaviest N=Z region
 - Super-allowed α decay
 - Enhanced proton-neutron correlation
 - Enhancement in E2 collectivity?
 - Role of proton-neutron pairing
 - Breaking of the isospin symmetry
 - 2. Aspect of a good playground for investigating the shell evolution
 - Sn (Z=50 magic) chains:
 - long stable isotope chain
 - proton-rich side: accessible with fusion reactions
 - neutron-rich side: accessible with fission or fragmentation

Recent experimental findings



I. G. Darby et al., Phys. Rev. Lett. 105, 162502 (2010).

Proton shell structure seen in 51Sb isotopes

reduction of the *ls* splitting for larger N?

shell evolution due to the tensor force



J. P. Schiffer et al., Phys. Rev. Lett. 92, 162501 (2004).



T. Otsuka et al., Phys. Rev. Lett. 95, 232502 (2005).

Contradicting experiment and picture

- (³He, d) reaction
 - much smaller C²S (≈ 0.5 for $11/2^{-}$)
 - Strong influence of correlation such as π(d_{5/2})*3⁻ has been suggested, for instance, by Sorlin et al.

this work

Shell-model calculations in the full $50 \le N(Z) \le 82$ space are carried out.

 Both shell evolution and correlation are taken into account.



O. Sorlin and M.-G. Porquet, Prog. Part. Nucl. Phys. 61, 602 (2008).

$V_{\rm MU}$: monopole-based universal interaction

- proposed for a universal shell evolution: two ingredients
 - 1. tensor force
 - The "bare" tensor force appears good, but it is not obvious whether it works as the effective interaction
 - From microscopic theory, the effective tensor force is almost irrelevant to the normalization procedures due to
 - short-range correlation
 - finite size of model space

Use of π+ρ tensor force is justified by "renormalization persistency."



T. Otsuka et al., Phys. Rev. Lett. 104, 012501 (2010); N. Tsunoda, T. Otsuka et al., Phys. Rev. C 84, 044322 (2011).

$V_{\rm MU}$: monopole-based universal interaction

- two ingredients
 - 2. central force
 - phenomenological simplicity for the tensor-subtracted force



proposed universal effective force:



Qualitative difference between central and tensor in terms of shell evolution



- tensor: spin dependence \rightarrow responsible for *l*·*s* term
- central: node dependence \rightarrow responsible for l^2 term

Example of success of V_{MU} : Is splitting of ⁴⁸Ca



Y. Utsuno et al., Phys. Rev. C 86, 051301(R) (2012).

Some details about the interaction

- Proton-neutron interaction
 - dominating the evolution in Sb isotopes
 - The overall strength of the central part of
 $V_{\rm MU}$ is tuned so as to reproduce the
 proton separation energies of Sb isotopes.
 - \rightarrow only one free parameter
- Neutron-neutron interaction
 - responsible for making "Sn core"
 - SNBG3 (n-n) interaction by Honma et al.
 - a semi-empirical interaction fitted to Sn isotopes



M. Honma et al., RIKEN Accel. Prog. Rep. (2012).

Proton single-particle energies



- lack of data on SPE on top of ¹⁰⁰Sn (except 1st excited state in ¹⁰¹Sn)
- reference states: ¹³³Sb (N=82)
 - single-particle(-like) states below 3 MeV because the 1st excited state in ¹³²Sn is as high as 4 MeV.
- Single-particle levels of ¹⁰¹Sb are thus a prediction.

Proton single-particle energies on top of ¹³²Sn core

level	7/2+	5/2+	3/2+	11/2-	1/2+
Ex. (MeV)	0	0.962	2.440	2.791	2.990

Experimental value except 1/2⁺ taken from systematics of W. J. Baldridge et al.

Evolution of the energy levels

- 11/2⁻₁ and 5/2⁺₁ levels
 measured from 7/2⁺₁
 - full shell-model results vs.
 estimate from effective
 single-particle energies
 (ESPE)
 - Non-monotonic evolution is reproduced well.
 - Most of the decrease of 11/2- level (2.25 MeV from N=82 to 62) is explained with the shell model, but is not with the ESPE.



Estimate from ESPE

- Definition of ESPE
 - For an assumed core |core>, ESPE of the orbit j is defined as

$$\tilde{\varepsilon}_{i} = \left\langle \text{core} \left| a_{i} H a_{i}^{\dagger} \right| \text{core} \right\rangle - \left\langle \text{core} \left| H \right| \text{core} \right\rangle$$

$$=\varepsilon_i + \sum_{j:occupied} v_{i,j} n_j$$

- Two ways of assumed core (different in terms of n_i)

- 1. naïve filling configuration (ESPE1)
- 2. calculated ground state of Sn isotopes (ESPE2)



Evolution of the 11/2⁻-7/2⁺ spacing

- Consider the change from N=64 (to 82): 2.215 MeV in expt.
 - ESPE1: no correlation
 - (6, 8, 0, 0, 0) for occupation in $(d_{5/2}, g_{7/2}, h_{11/2}, s_{1/2}, d_{3/2})$ $n(j_{<})-n(j_{>})=2$
 - tensor only: 1.93 MeV
 - central + ls + tensor: 1.93 MeV
 - ESPE2: n-n correlation only
 - (4.76, 5.91, 1.74, 0.69, 0.91) $n(j_{<})-n(j_{>})=0.32$
 - tensor only: 1.21 MeV
 - central + ls + tensor: 1.38 MeV
 - shell-model calculation: full correlation
 - central + ls + tensor: 1.89 MeV





T. Otsuka et al., Phys. Rev. Lett. 95, 232502 (2005).

Single-particle strength



- Relative quantities are very stable from N=62 to 74.
- the absolute strength: should be large or small?

Evolution of μ from N=70 to 82

- μ (magnetic moment): very good probe to examine the single-particle property
 - Configuration mixing should change μ from the singleparticle value.
 - Particle-vibration coupling

$$\mu(j) = \left(1 - \frac{3}{j(j+1)} \frac{\alpha^2}{1+\alpha^2}\right) \mu_{\rm sp}(j) + \frac{3}{2(j+1)} \frac{\alpha^2}{1+\alpha^2} \mu(2_1^+),$$

 Excellent agreement with experiment indicating the increase of mixing



effective nucleon *g* factor adopted:

isovector shift $g_i(IV)=0.1$, spin quenching factor 0.6

 $\Rightarrow \mu(^{133}Sb)=2.97$ (calc.) vs. 3.00(1) (expt.)

How about μ of $11/2^-$?

- Only ^{115,117}Sb are available experimentally.
- Mixing almost always reduces μ of 11/2⁻ because single-particle value 7.18 is very large (due to j_> with high /).
 - For instance, μ of $\pi(d_{5/2})^*3^-$ is 3.27 using the shell-model wave function for 3^- .
- Calculated value is reduced from the single-particle value to some extent and agrees well with experiment:
 Considerable configuration mixing should occur.

μ of the 11/2 $^{\scriptscriptstyle 2}$ states

	Expt.	Calc.
¹¹⁵ Sb	5.53(8)	5.53
¹¹⁷ Sb	5.35(9)	5.63

Expt.: taken from ENSDF

effective nucleon g factor adopted: isovector shift $g_i(IV)=0.1$, spin quenching factor 0.6

Possible experiment at JAEA Tandem

- 5/2⁺: goes up or down toward ¹⁰¹Sb?
 - SPE due to the Gogny D1S does not show the turning.
- Previous α decay experiment of ¹⁰⁹I (C. Mazzaocchi et al., Phys. Rev. Lett. 98, 212501 (2007).) did not observe a γ ray.
 - severe competition with proton decay?



Proton SPE due to Gogny DIS

¹⁰⁰Sn

¹⁰²Sn

¹⁰⁴Sn



Summary

- The evolution of the low-lying energy levels in Sb isotopes is studied with large-scale shell-model calculations.
 - shell evolution vs. correlation
 - universal treatment of the p-n interaction using $V_{\rm MU}$: most important part for the shell evolution
- Experimental energy levels of 11/2⁻ and 5/2⁺ measured from 7/2⁺ are reproduced satisfactorily.
 - The evolution of SPE mainly due to the tensor force accounts for them partially, but the correlation is important from a quantitative point of view.
 - The spectroscopic factors around the mid-shell are almost half a singleparticle limit, which is supported by the evolution of the magnetic moment.
 - It is expected that predicted levels are measured at JAEA Tandem under the collaboration of JAEA, York, and Oak Ridge.